

# CENTER FOR LUNAR SCIENCE AND EXPLORATION

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NASA Johnson Space Center  
Astromaterials Research & Exploration Science

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OF ARIZONA

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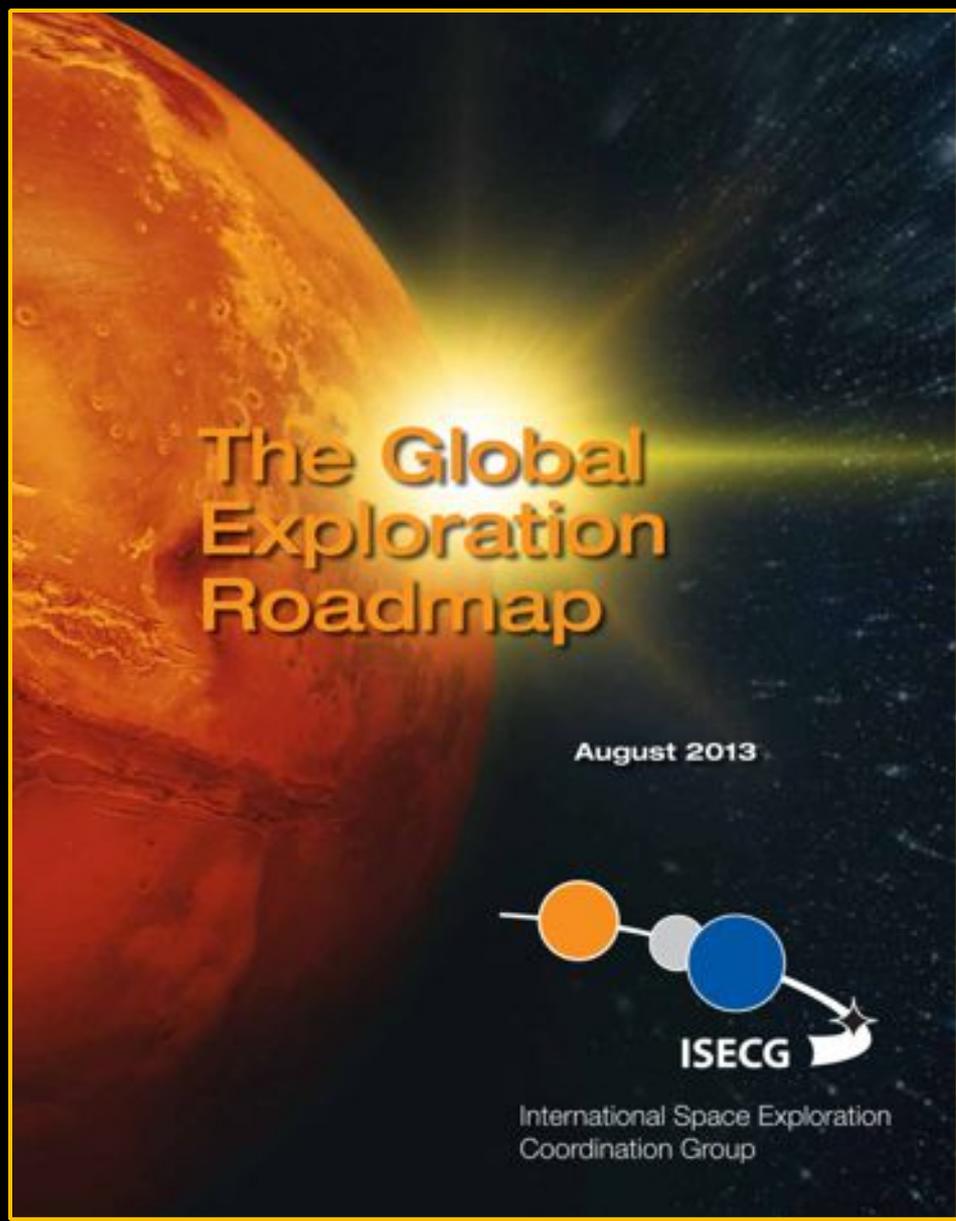
UNIVERSITY OF  
MARYLAND

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## Science Questions & How They Might Be Addressed By the Presence of Humans/ Human Support Structure

David A. Kring

Lunar and Planetary Institute



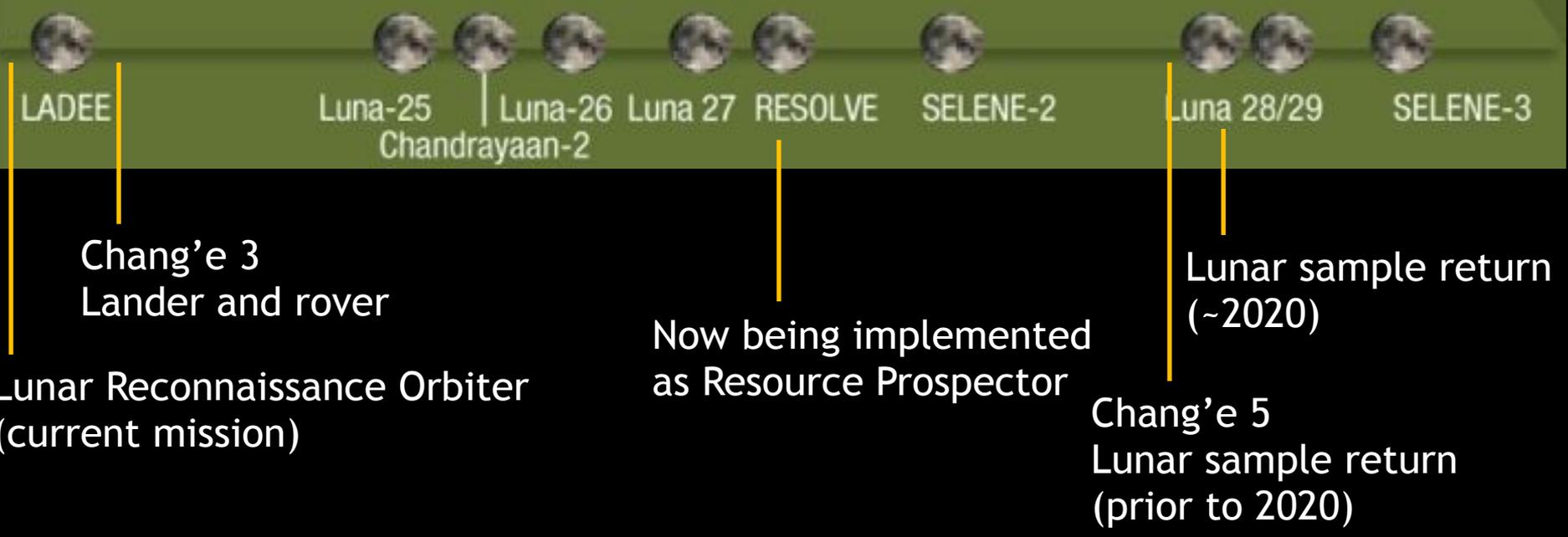
## Roadmap for Human Exploration

- Outlines a plan that extends human exploration beyond low-Earth orbit (LEO)
- Includes multiple destinations (the Moon, asteroids, and eventually Mars)
- Highlights the need for a robotic program that
  - Serves as precursor explorers, then as
  - A parallel mission element partner,
  - Before we have a fully developed human exploration program

## Elements of a Lunar Robotic Program

- It is essential that we restore the capability of lunar surface operations and
- Sample return

### Robotic Missions to Discover and Prepare



## NASA's Resource Prospector Robotic Lunar Surface Element



### RESOLVE (Regolith & Environment Science and Oxygen and Lunar Volatile Extraction)



#### Sample Acquisition –

##### Auger/Core Drill [CSA provided]

- Complete core down to 1 m; Auger to 0.5 m
- Minimal/no volatile loss
- Low mass/power (<25 kg)
- Wide variation in regolith/rock/ice characteristics for penetration and sample collection
- Wide temperature variation from surface to depth (300K to <100K)

#### Sample Evaluation –

##### Near Infrared Spectrometer (NIR)

- Low mass/low power for flight
- Mineral characterization and ice/water detection before volatile processing
- Controlled illumination source

#### Resource Localization –

##### Neutron Spectrometer (NS)

- Low mass/low power for flight
- Water-equivalent hydrogen  $\geq 0.5$  wt% down to 1 meter depth at 0.1 m/s roving speed

#### Volatile Content/Oxygen Extraction –

##### Oxygen & Volatile Extraction Node (OVEN)

- Temperature range of <100K to 900K
- 50 operations nominal
- Fast operations for short duration missions
- Process 30 to 60 gm of sample per operation (Order of magnitude greater than TEGA & SAM)

#### Volatile Content Evaluation –

##### Lunar Advanced Volatile Analysis (LAVA)

- Fast analysis, complete GC-MS analysis in under 2 minutes
- Measure water content of regolith at 0.5% (weight) or greater
- Characterize volatiles of interest below 70 AMU

#### Operation Control –

##### Flight Avionics [CSA/NASA]

- Space-rated microprocessor

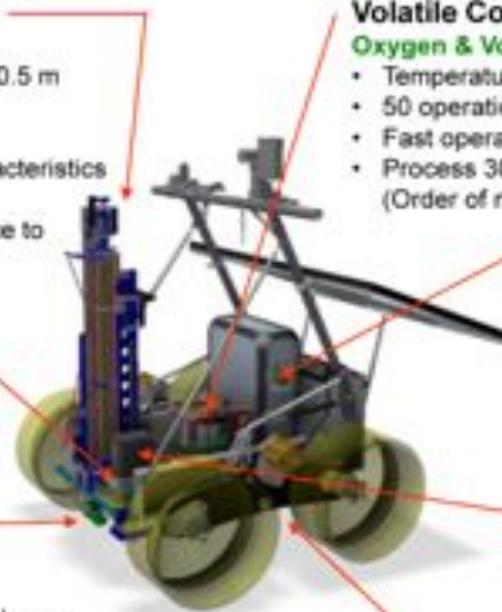
#### Surface Mobility/Operation [CSA mobility platform]

Rover nicknamed "Artemis Jr."

- Low mass/large payload capability
- Driving and situation awareness, stereo-cameras
- Autonomous navigation using stereo-cameras and sensors
- NASA contributions likely for communications and thermal management

#### RESOLVE Instrument Suite Specifications

- Nom. Mission Life = 10+ Cores, 12+ days
- Mass = 60-70 kg
- Dimensions = w/o rover: 68.5 x 112 x 1200 cm
- Ave. Power: 200 W

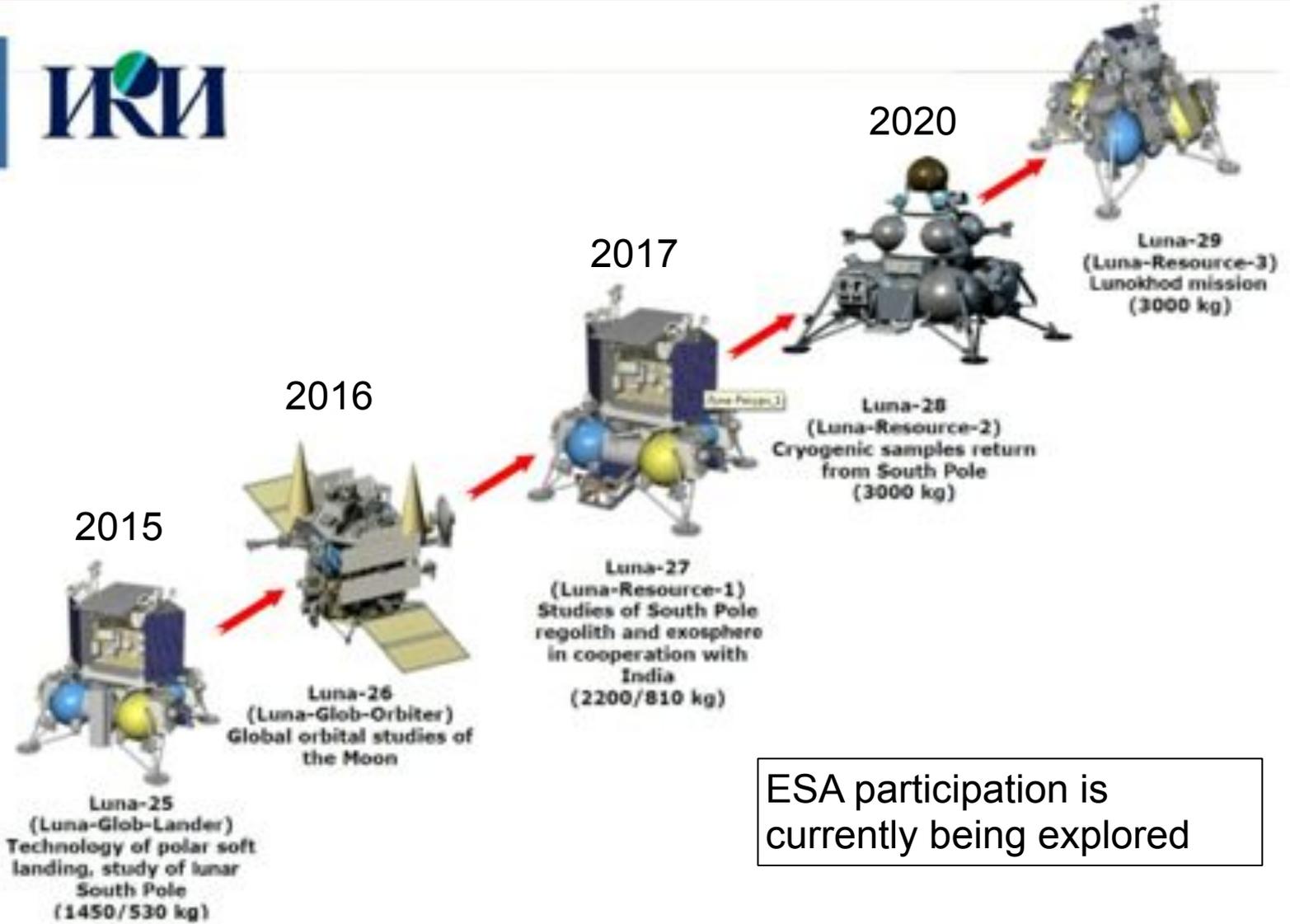




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## A Roscosmos-led Series of Robotic Lunar Surface Elements



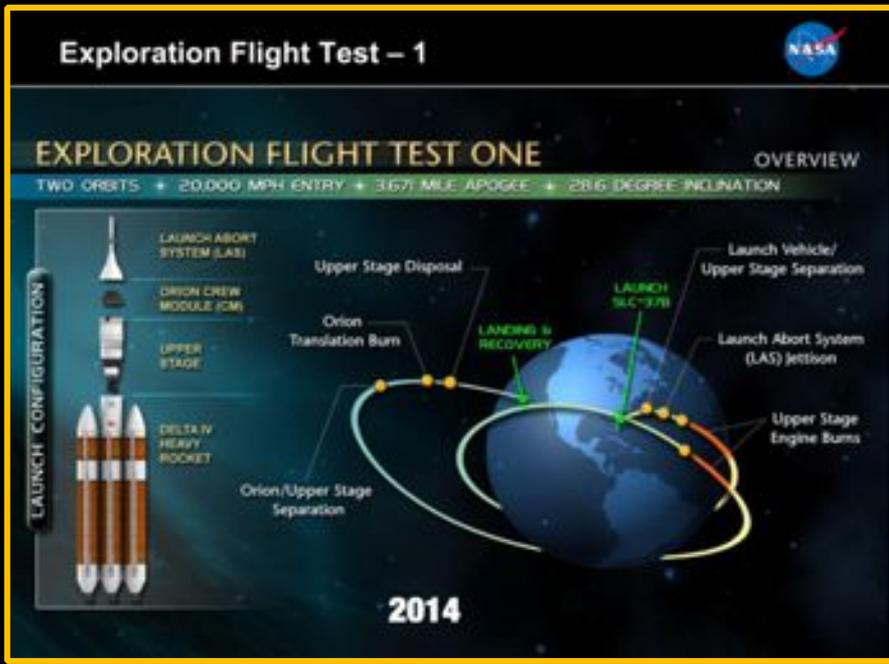
ESA participation is currently being explored



# CENTER FOR LUNAR SCIENCE AND EXPLORATION

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## EXPLORATION – DEVELOPMENT OF ORION AND SLS VEHICLES



EFT-1 (December 2014)



# CENTER FOR LUNAR SCIENCE AND EXPLORATION

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## EXPLORATION – DEVELOPMENT OF ORION AND SLS VEHICLES

### Exploration Flight Test – 1



### EXPLORATION FLIGHT TEST ONE

OVERVIEW

TWO ORBITS + 20,000 MPH ENTRY + 3,671 MILE APOGEE + 29.6 DEGREE INCLINATION



EFT-1 (December 2014)

EM-1 and EM-2 (2017 and 2021)

### ESD Mission Overview



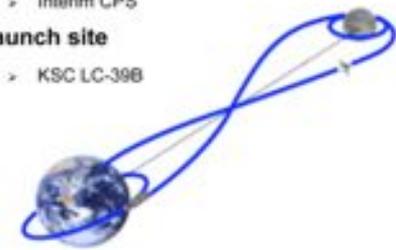
**Exploration Mission One (EM-1)**  
First Uncrewed BEO Flight  
2017

- **Mission objectives**
  - > Demonstrate integrated spacecraft systems performance prior to crewed flight
  - > Demonstrate high speed entry (~11 km/s) and TPS prior to crewed flight
- **Mission description**
  - > Un-crewed circumlunar flight – free return trajectory
  - > Mission duration ~7 days
- **Spacecraft configuration**
  - > Orion Uncrewed
- **Launch vehicle configuration**
  - > SLS Block 1, 5-segment RSRMV, 4 RS-25, 70mt
  - > Interim CPS
- **Launch site**
  - > KSC LC-39B



**Exploration Mission Two (EM-2)**  
First Crewed BEO Flight  
2021

- **Mission objectives**
  - > Demonstrate crewed flight beyond LEO
- **Mission description**
  - > Crewed lunar orbit-capable, or other destinations
  - > Mission duration 10-14 days
- **Spacecraft configuration**
  - > Orion Crewed
- **Launch vehicle configuration**
  - > SLS Block, 5-segment RSRMV, 4 RS-25, 70mt
  - > Interim CPS
- **Launch site**
  - > KSC LC-39B



## EXPLORATION – DEVELOPMENT OF ORION AND SLS VEHICLES

### Exploration Flight Test – 1

#### EXPLORATION FLIGHT TEST ONE

OVERVIEW

TWO ORBITS + 20,000 MPH ENTRY + 3,671 MILE APOGEE + 28.6 DEGREE INCLINATION



### ESD Mission Overview

#### Exploration Mission One (EM-1) First Uncrewed BEO Flight 2017

##### • Mission objectives

- Demonstrate integrated spacecraft systems performance prior to crewed flight
- Demonstrate high speed entry (~11 km/s) and TPS prior to crewed flight

##### • Mission description

- Un-crewed circumlunar flight – free return trajectory
- Mission duration ~7 days

##### • Spacecraft configuration

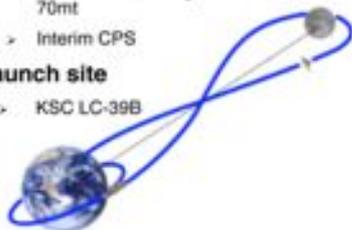
- Orion Uncrewed

##### • Launch vehicle configuration

- SLS Block 1, 5-segment RSRMV, 4 RS-25, 70mt
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##### • Launch site

- KSC LC-39B



EFT-1 (December 2014)

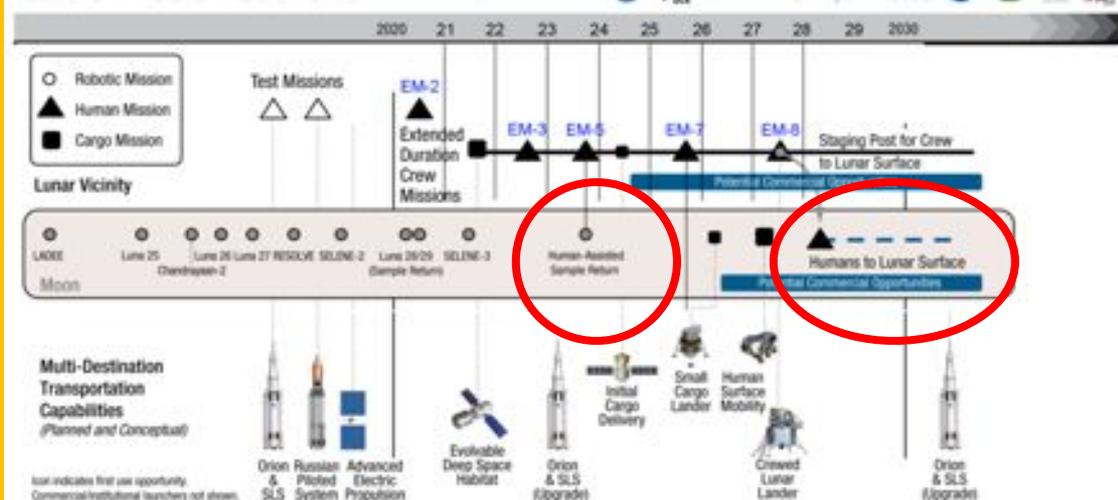
EM-1 and EM-2 (2017 and 2021)

And, within the framework of the Global Exploration Roadmap:

- Human-assisted sample return
- Humans to the lunar surface

➢ Demonstrate crewed flight beyond LEO

### ISECG Mission Scenario



# NASA Human Spaceflight Architecture Team (M. Lupisella and M. R. Bobskill, 2012)

## Preliminary Mission Sequence



### Mission Summary

- Crew visits Deep Space Facility (DSF) located @ E-M Lagrange point. Crew tests & demonstrates future exploration systems & operations, controlling assets on the lunar surface, performing lunar surface observations, and other deep space science tasks. DSF could be derivative of ETM and moved between E-M L1 & L2 points and be visited multiple times.
- Uses ETM and MPCV (first MPCV/SLS mission beyond test flights) for crew pressurized volume
- First crew arrives for first DSF mission, later crews bring further infrastructure, stay for longer durations
- Station-keeping (with ACS, RCS?)
- NEA stack assembly in situ?

### Mission Benefits

- Develop habitation capabilities & reduce risk for future exploration missions
- Enhance lunar & space science: e.g., survey farside, control robots on surface, perform surface assembly (e.g. farside radio telescope)
- Perform exploration research & technology ops: crew+robot autonomous ops, long delay comm, advanced EVA systems, measurements (e.g., radiation shielding)
- Demonstrate deep space assembly
- ETM serves as foothold in deep space

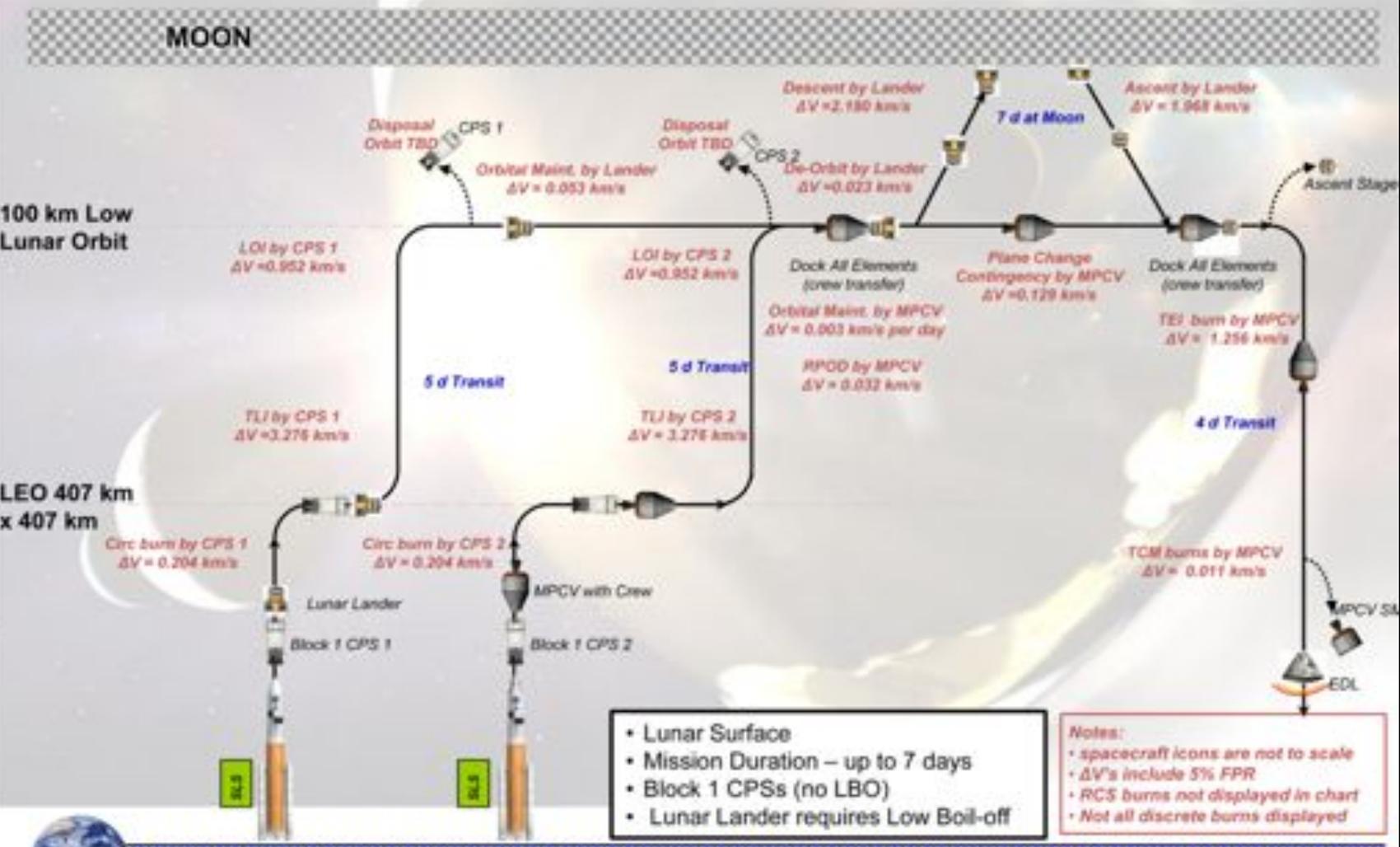


# NASA Human Spaceflight Architecture Team (Connolly et al. , 2012)

## Lunar Destination Activities



### Lunar Sortie DRM



- Lunar Surface
- Mission Duration – up to 7 days
- Block 1 CPSs (no LBO)
- Lunar Lander requires Low Boil-off

**EARTH**



## The Missing Element

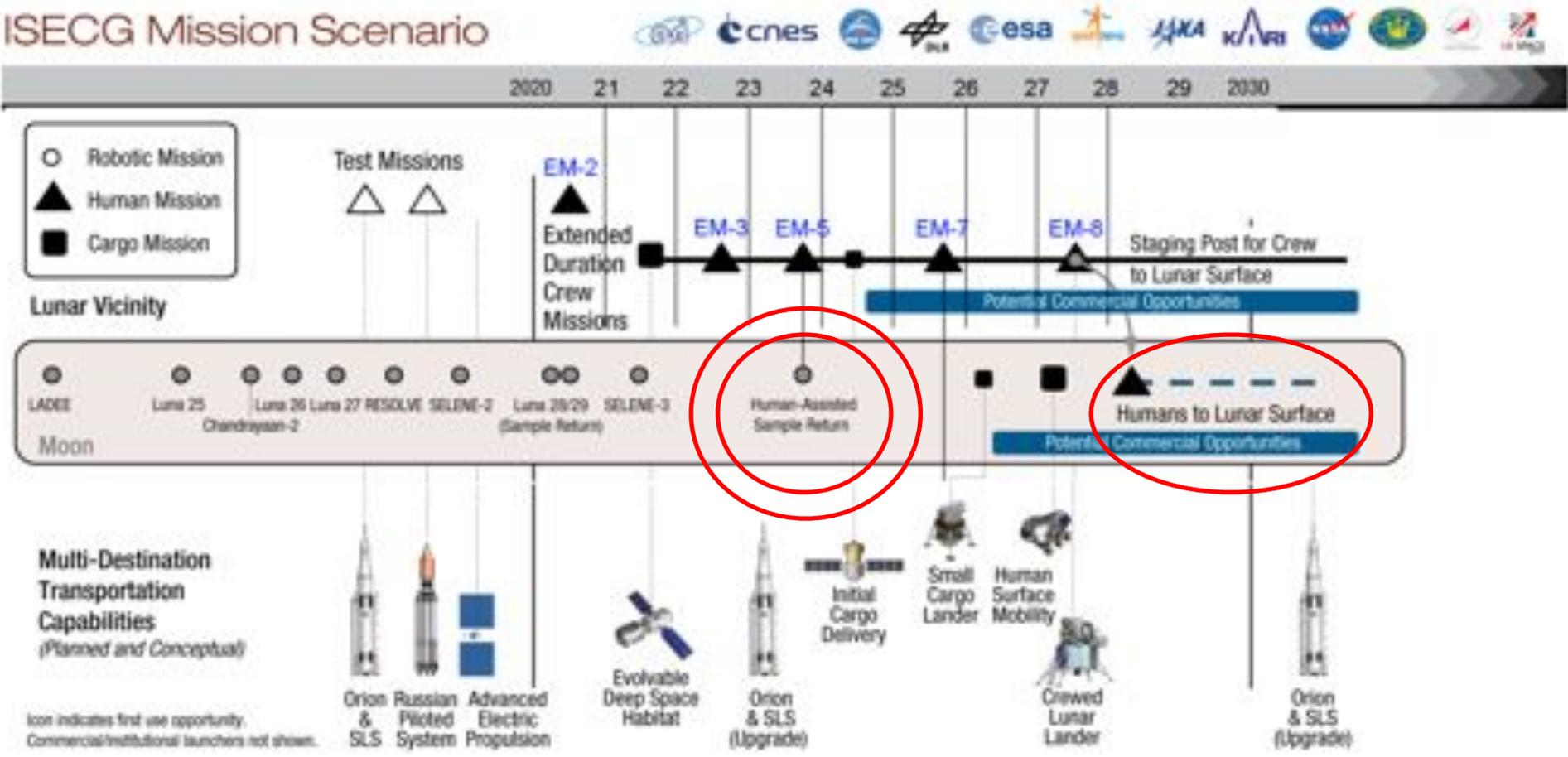
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- Human-rated lunar lander (European concept pictured)
- Until this capability is developed, we will be limited to missions with crew in cis-lunar space and robotic lunar surface components

## Developing the Human Exploration Elements

- NASA's SLS and Orion vehicles
- ESA service module

### ISECG Mission Scenario



Detail of illustration from the GER (2013) with small modifications



*The Scientific Context for*  
**EXPLORATION**  
*of the*  
**MOON**

NATIONAL RESEARCH COUNCIL  
OF THE NATIONAL ACADEMIES

In 2007,

The National Research Council published a report called *The Scientific Context for Exploration of the Moon*, which provided NASA with scientific guidance for an enhanced exploration program that would provide global access to the lunar surface through an integrated robotic and human architecture.

The report outlined 3 major hypotheses, identified 8 science concepts, and, within those concepts, it

Identified 35 specific investigations

Importantly, the report also prioritized those investigations



*The Scientific Context for*  
**EXPLORATION**  
*of the*  
**MOON**

NATIONAL RESEARCH COUNCIL  
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The “BIG” ideas to be explored:

- Giant impact hypothesis for the origin of the Moon



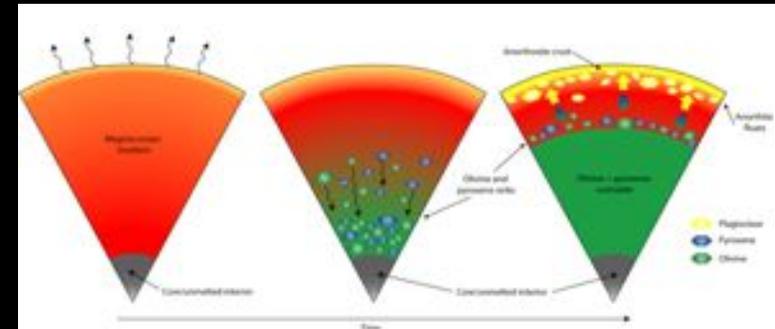


*The Scientific Context for*  
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*of the*  
**MOON**

NATIONAL RESEARCH COUNCIL  
OF THE NATIONAL ACADEMIES

The “BIG” ideas to be explored:

- Giant impact hypothesis for the origin of the Moon
- Lunar magma ocean hypothesis and fundamental principles of planetary differentiation





*The Scientific Context for*  
**EXPLORATION**  
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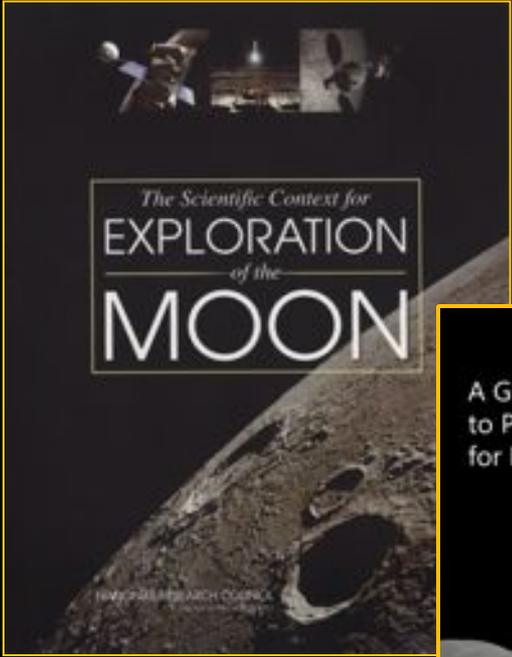
NATIONAL RESEARCH COUNCIL  
OF THE NATIONAL ACADEMIES

The “BIG” ideas to be explored:

- Giant impact hypothesis for the origin of the Moon
- Lunar magma ocean hypothesis and fundamental principles of planetary differentiation
- Lunar cataclysm and inner solar system cataclysm hypotheses



## Where on the Moon can these objectives be addressed?



Input



Output

Guided, in part, by the 2007 National Research Council report called *The Scientific Context for Exploration of the Moon*,

We conducted a six-year series of summer studies to identify suitable landing sites on the Moon.

Stacking maps of the locations where each investigation could be addressed, we identified sites where a particularly large range of objectives could be addressed at the same time.

A Global Lunar Landing Site Study  
to Provide the Scientific Context  
for Exploration of the Moon

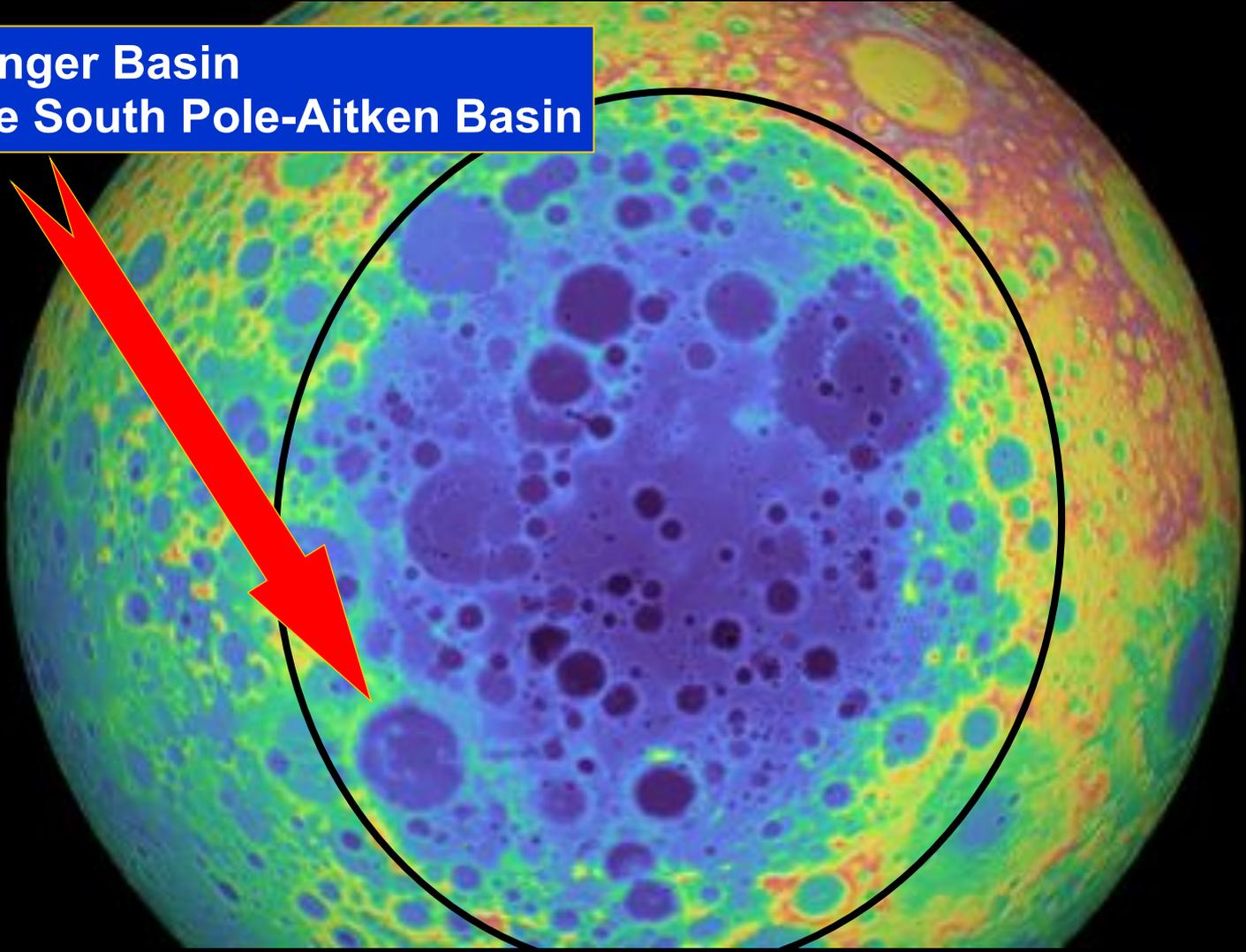


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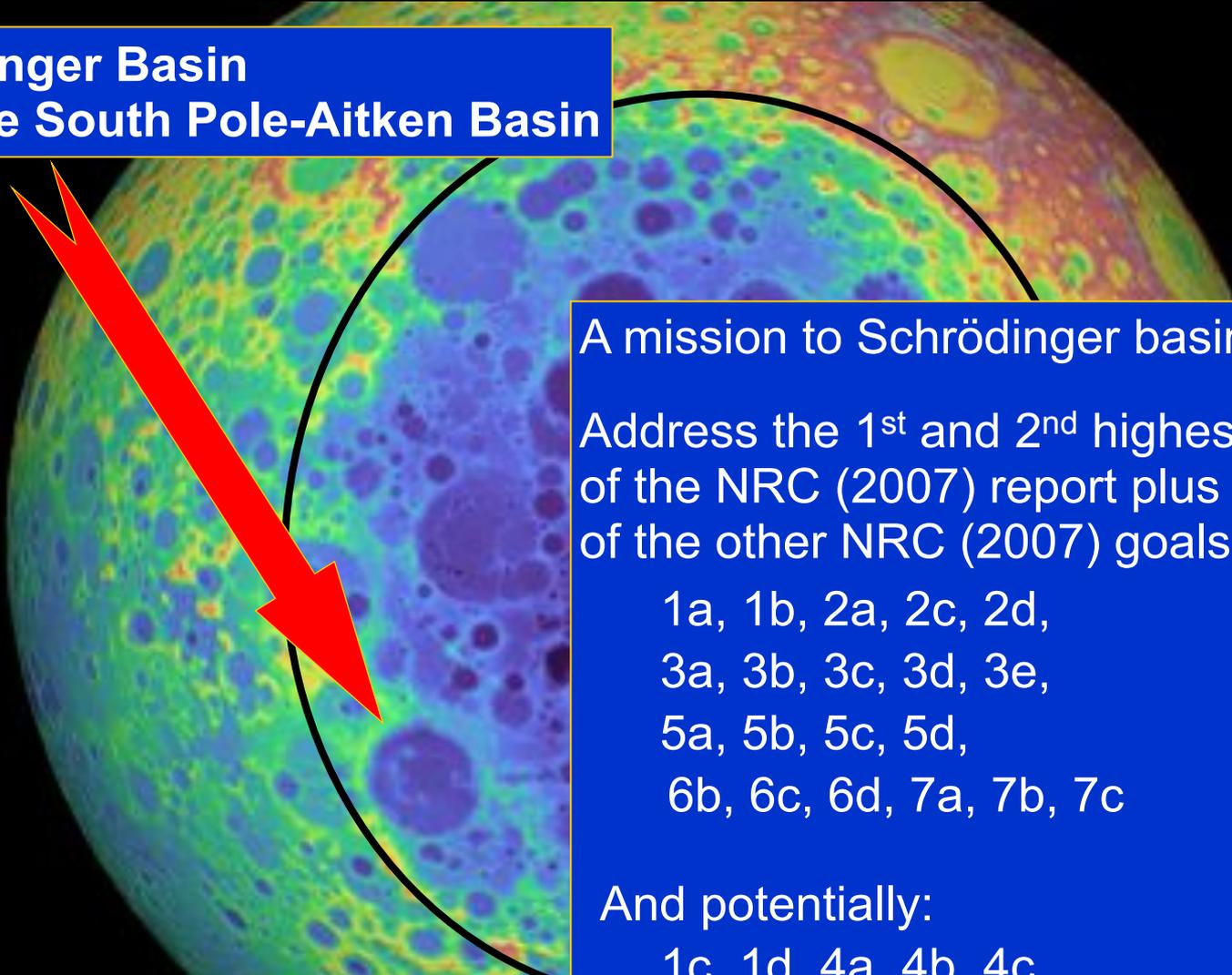
## Some highlights

- Schrödinger basin on the lunar far side, within the South Pole-Aitken basin, is the location where the largest range of objectives can be addressed.
- For studies of polar volatiles, Amundsen crater may be a better target than Shackleton crater.
- Most of the NRC (2007) objectives can be addressed within the South Pole-Aitken basin on the lunar far side,
- But to truly resolve all of the NRC (2007) objectives, global access to the Moon is required

**Schrödinger Basin  
w/i the South Pole-Aitken Basin**



## Schrödinger Basin w/i the South Pole-Aitken Basin



A mission to Schrödinger basin can:

Address the 1<sup>st</sup> and 2<sup>nd</sup> highest priorities of the NRC (2007) report plus many more of the other NRC (2007) goals:

- 1a, 1b, 2a, 2c, 2d,
- 3a, 3b, 3c, 3d, 3e,
- 5a, 5b, 5c, 5d,
- 6b, 6c, 6d, 7a, 7b, 7c

And potentially:

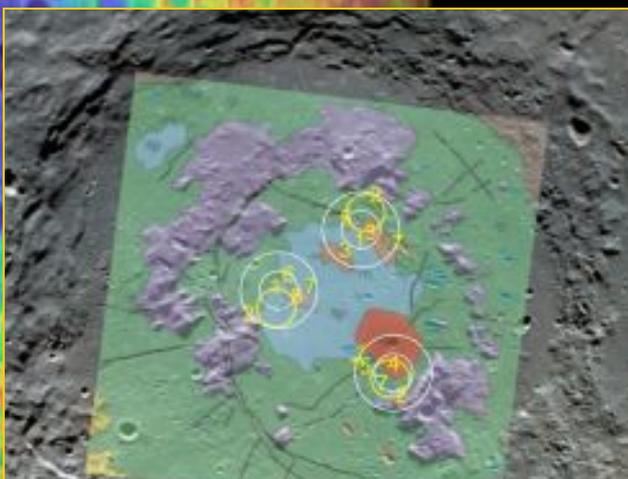
- 1c, 1d, 4a, 4b, 4c

## Schrödinger Basin w/i the South Pole-Aitken Basin



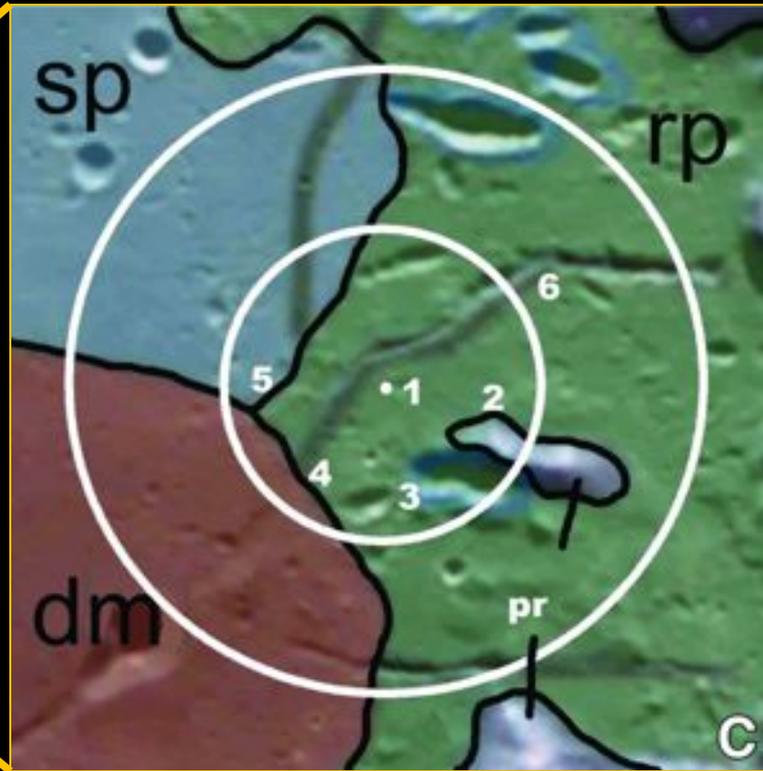
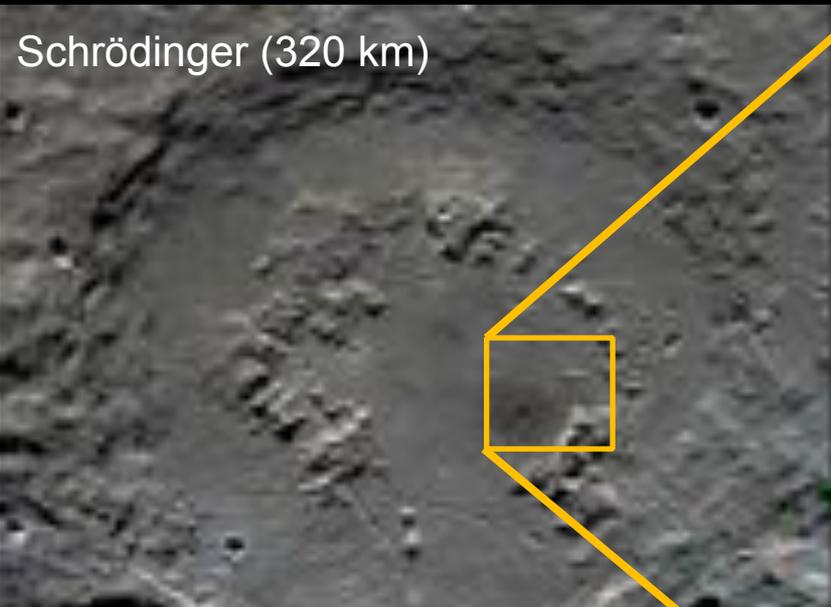
Schrödinger (320 km)

For those reasons, we have focused a lot of attention on Schrödinger basin. It is a very good target for future robotic and human exploration.



**Landing Site Study**  
O'Sullivan et al. (GSA SP 477, 2011)  
See also Bunte et al. (GSA SP 483, 2011)

## Schrödinger Basin within South Pole-Aitken Basin



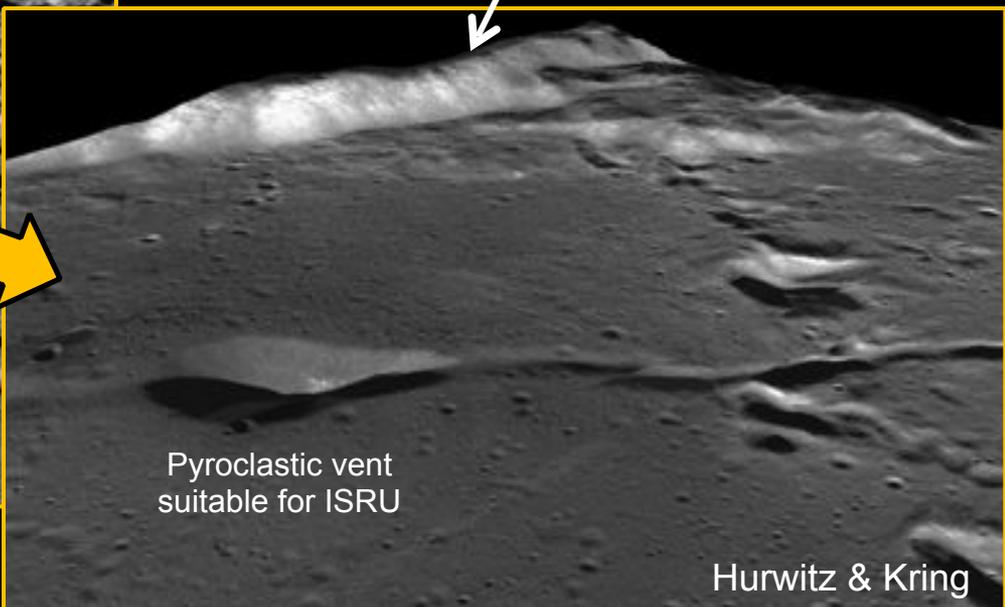
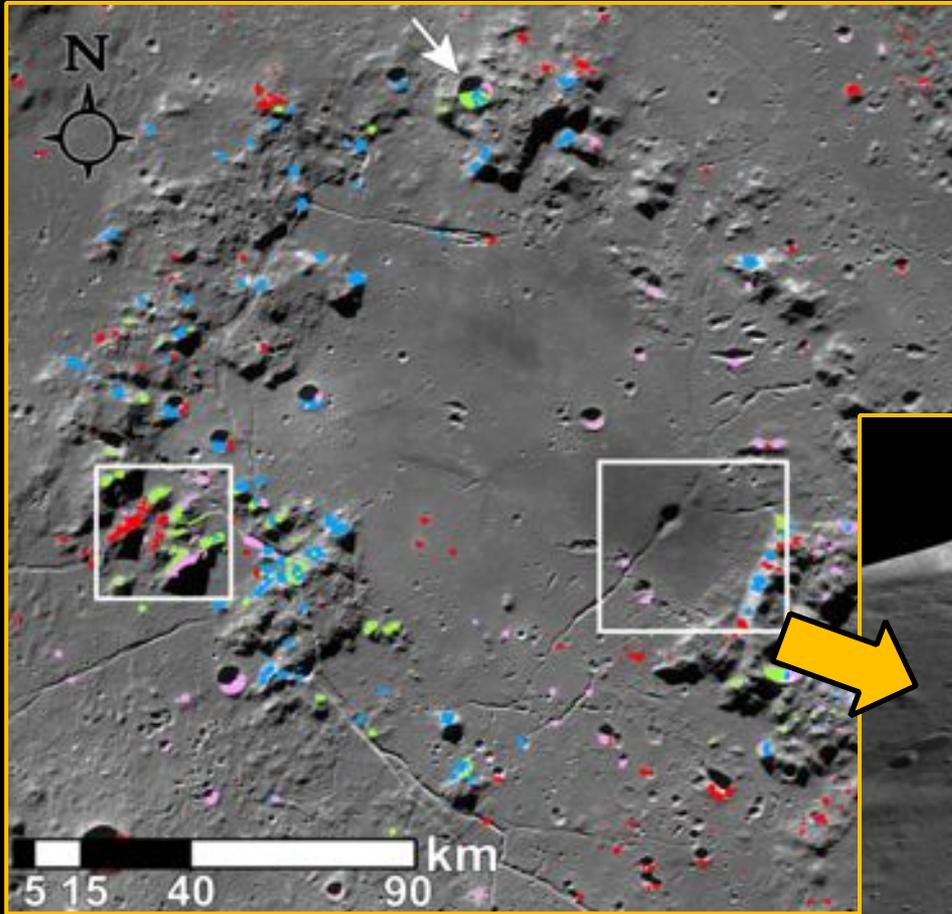
- Sta 1 = impact melt breccia
- Sta 2 = peak ring material
- Sta 3 = Antoniadi secondary crater
- Sta 4 = pyroclastic deposit
- Sta 5 = central melt sheet
- Sta 6 = deep fracture

O'Sullivan et al. (2011)

## Schrödinger Basin w/i the South Pole-Aitken Basin

Detailed studies by:  
Kramer, Kring, Nahm, & Pieters (Icarus 2013)  
Kumar et al. (JGR 2013)  
Chandnani et al. (LPSC 2013)

Using M<sup>3</sup> data, LOLA data, and LROC data.



Peak ring exposures of  
anorthositic, noritic, and troctolitic rocks

Pyroclastic vent  
suitable for ISRU

A Global Lunar Landing Site Study  
to Provide the Scientific Context  
for Exploration of the Moon

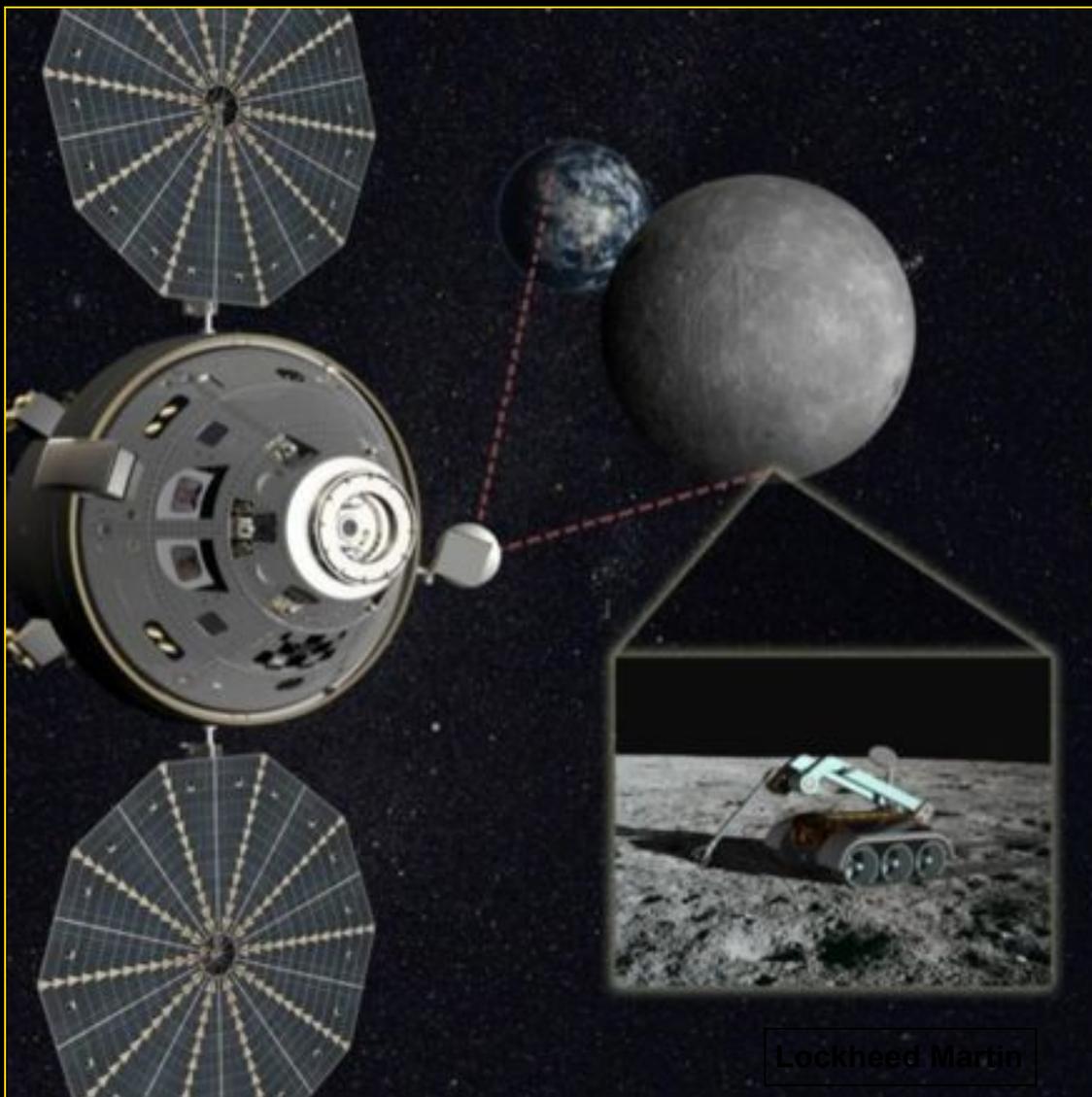


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## Experience suggests

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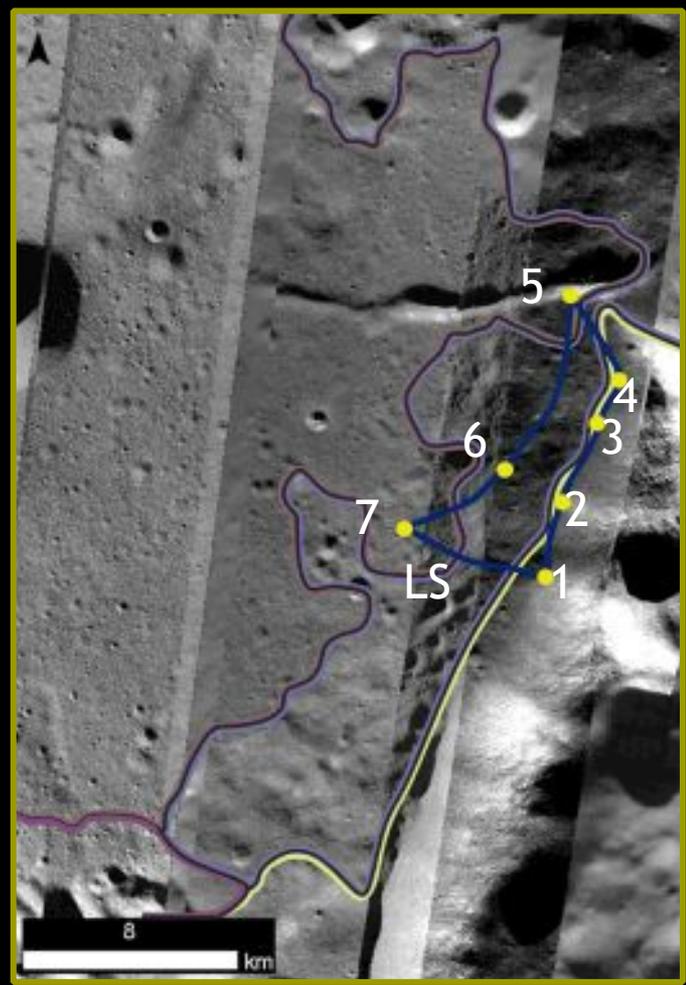
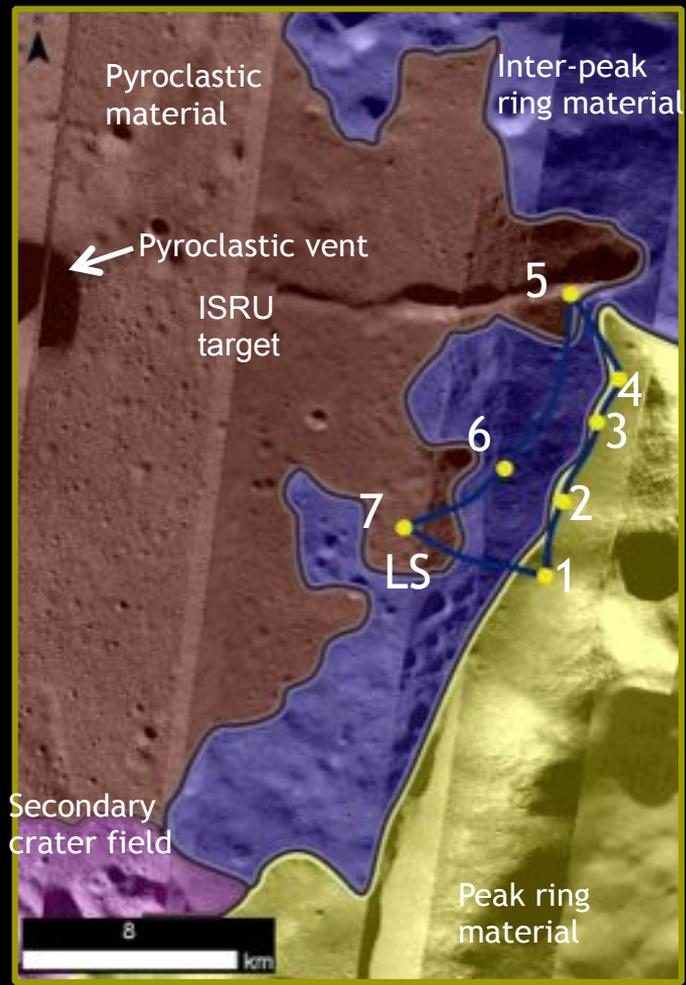
- **The best results would be obtained by a trained crew on the lunar surface**
- **If crew cannot be delivered to the lunar surface, then significant progress can be made robotically or with an integrated robotic and human architecture (e.g., deploying crew to Earth-Moon L2 above the lunar far side in Orion)**



## Re-examining the details:

- Our previous landing site study of **Schrödinger Basin** assumed crew were landing.
- In an integrated robotic and human exploration program that is consistent with the multi-agency **Global Exploration Roadmap**, we re-evaluated the landing site and stations for a robotic surface asset.

## EXAMPLE ROBOTIC TRAVERSE (SITE C)



**SITE C**  
28.8 km, 1 km/hr  
13.5 days (total traverse time)

Addresses NRC (2007)

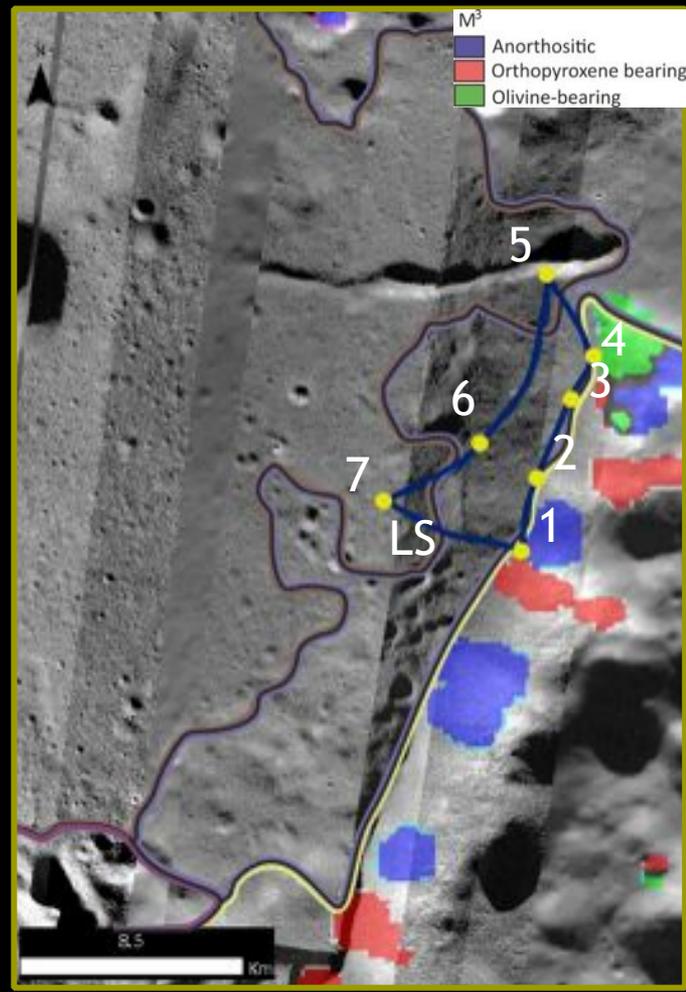
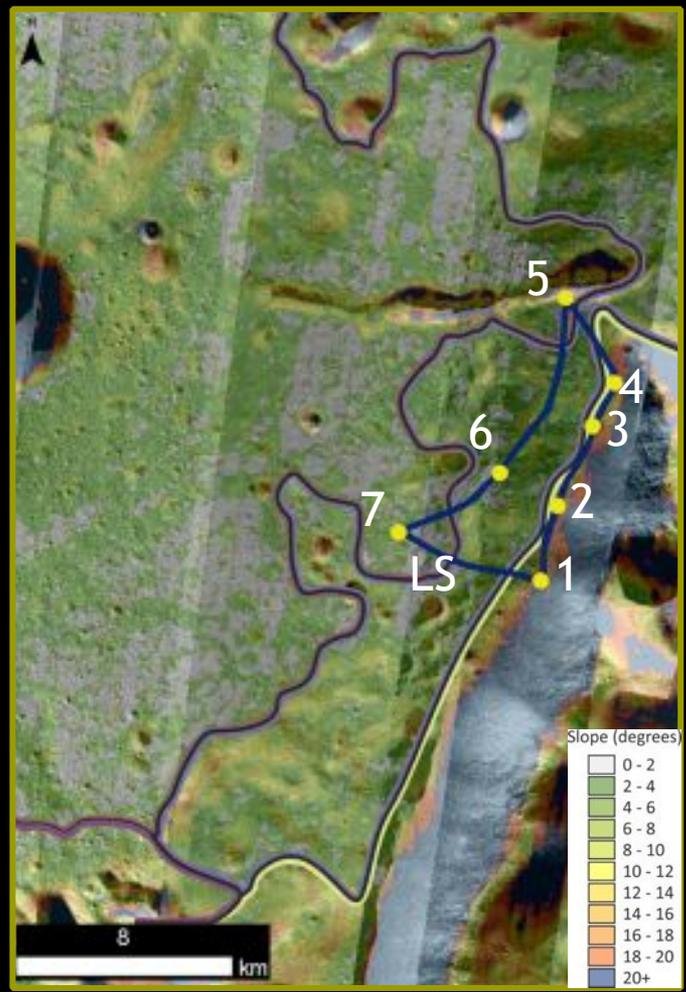
priority:

- Station 1: 2, 3, 7
- Station 2: 2, 3, 7
- Station 3: 2, 3, 7
- Station 4: 2, 3, 7
- Station 5: 2, 3, 5, 7
- Station 6: 1, 3, 6, 7
- Station 7: 3, 5, 6, 7

Plus ISRU studies in the vicinity of the pyroclastic vent

Gullickson et al. (2014)

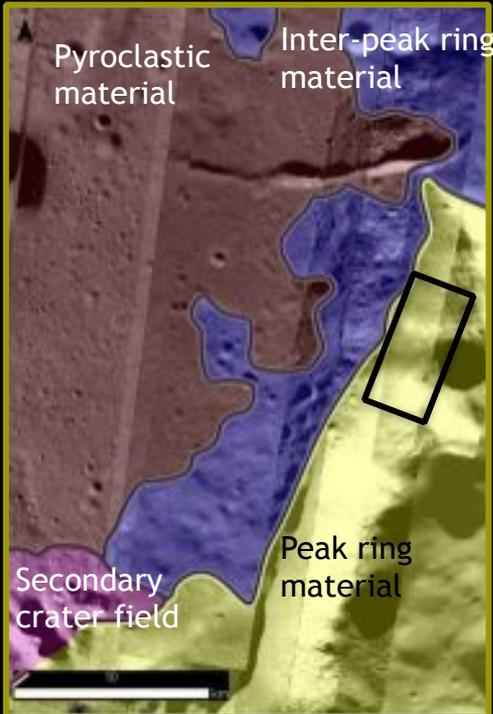
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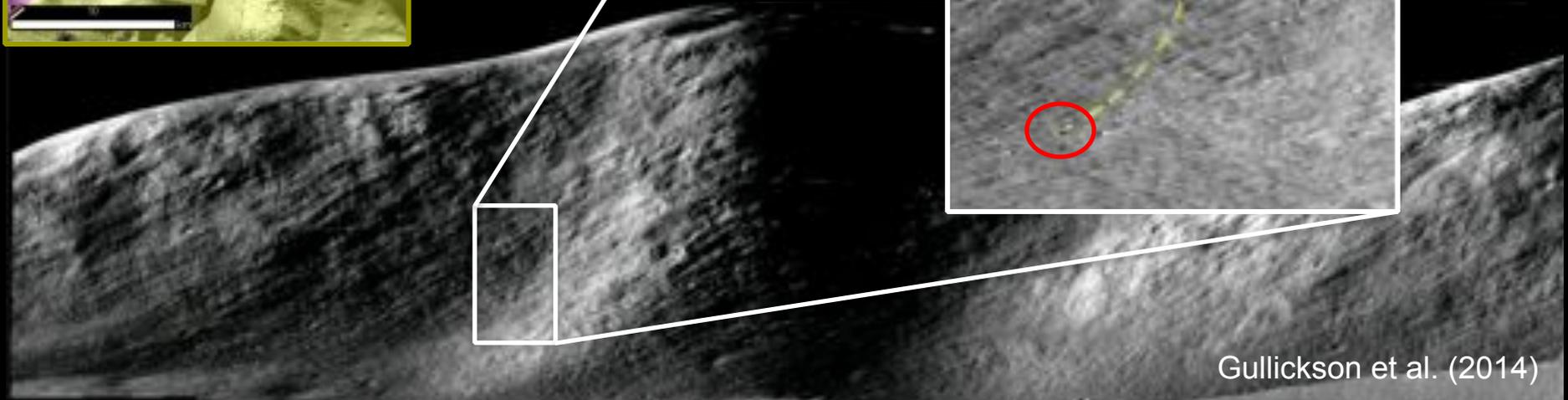
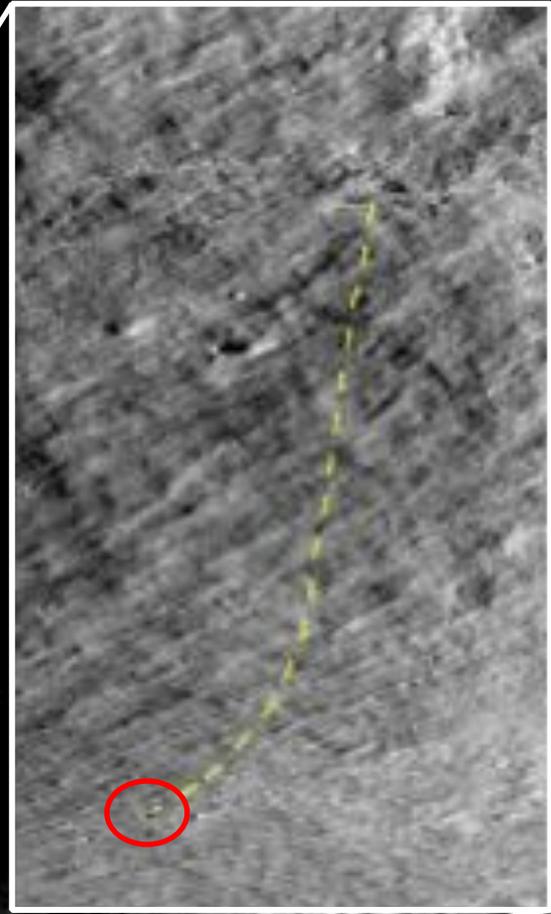
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Station 6: 1, 3, 6, 7  
Station 7: 3, 5, 6, 7

Plus ISRU studies in the vicinity of the pyroclastic vent  
Gullickson et al. (2014)



These studies have even identified the specific rocks that should be sampled





## EXAMPLE ROBOTIC TRAVERSE (SITE C)

**SITE C**  
28.8 km, 1 km/hr  
13.5 days

Average slope: 6.1°  
Maximum slope: 15.8°



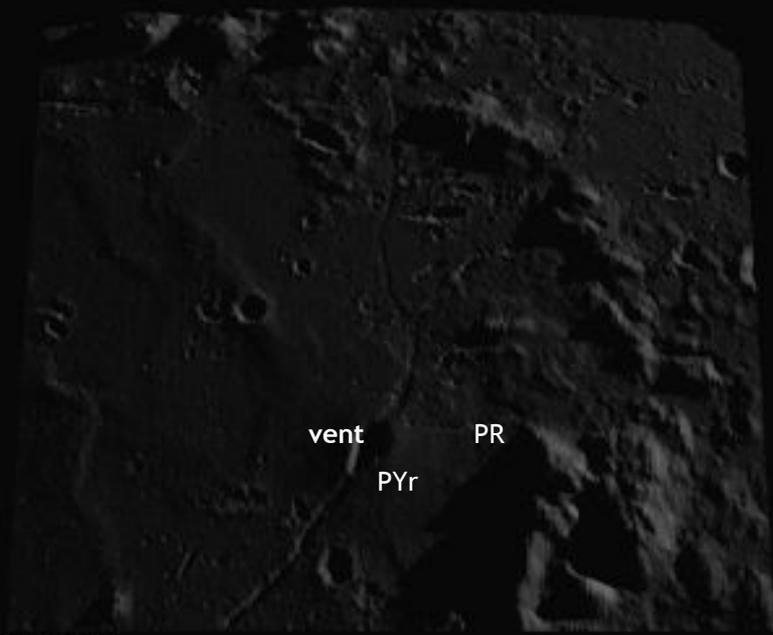
Where samples can be loaded into the ascent vehicle for return to Earth

Gullickson et al. (2014)



## ILLUMINATION -SITE C

- Mission planned 2021
- Optimum period of sunlight = 6<sup>th</sup> August 2021 - 19<sup>th</sup> August 2021

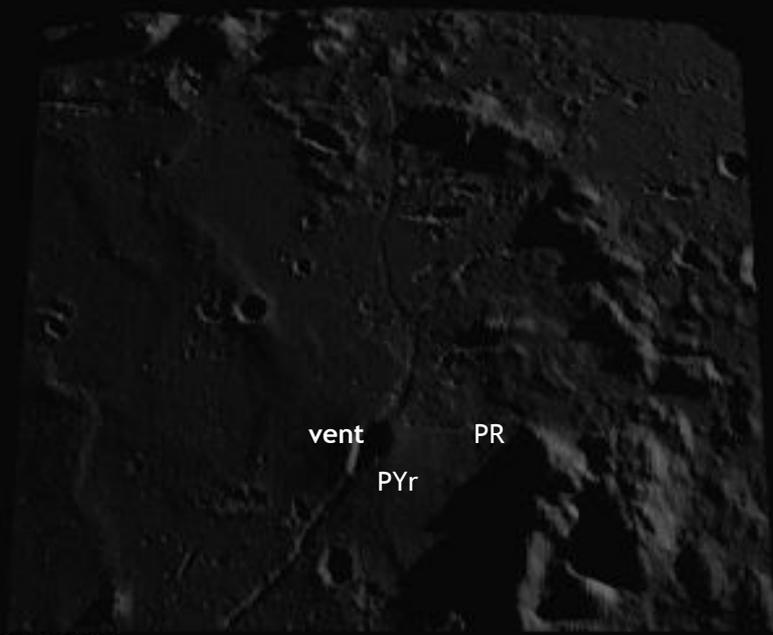


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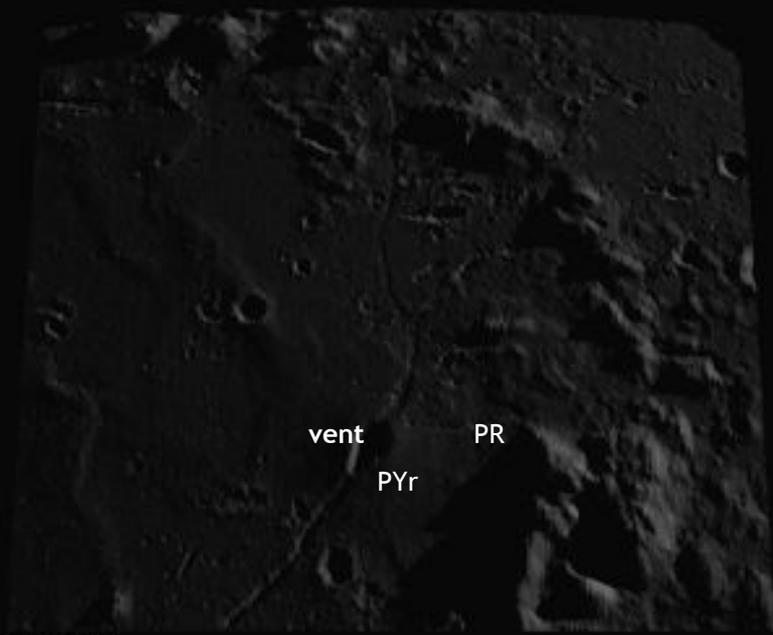


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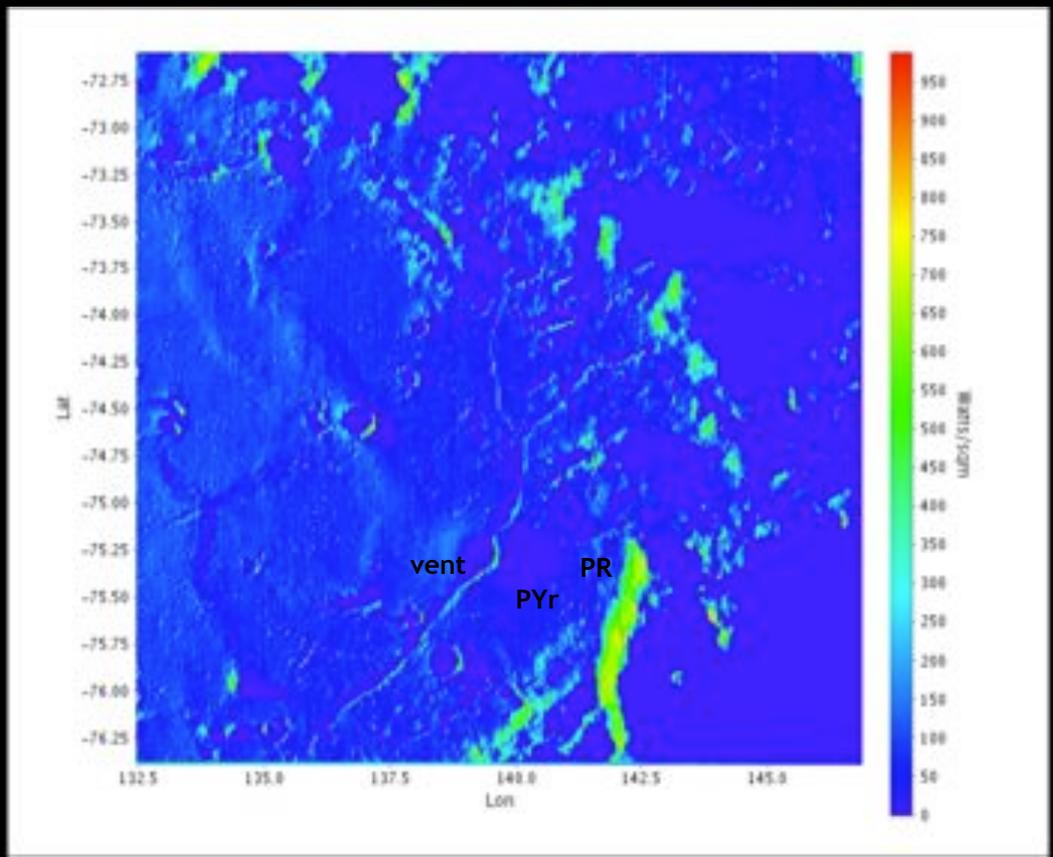


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## SOLAR IRRADIANCE - SITE C

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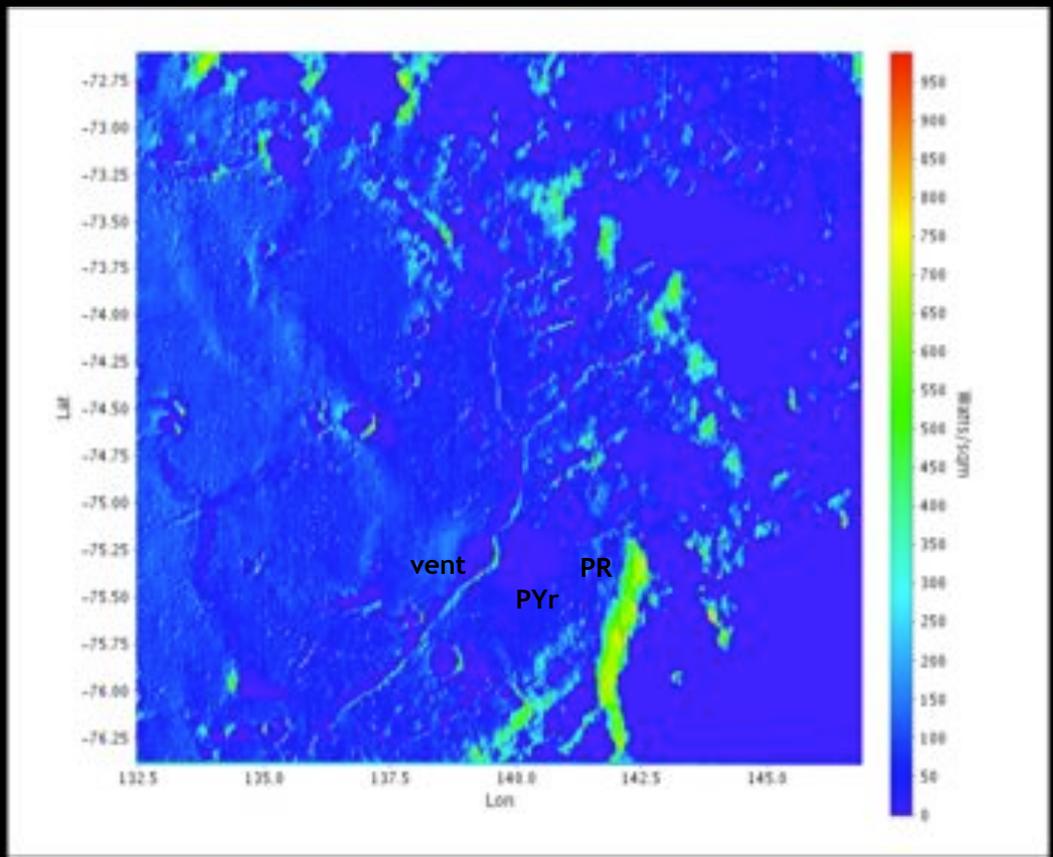
Incidence  
sunlight  
(i.e., sunlight  
power on  
surface)



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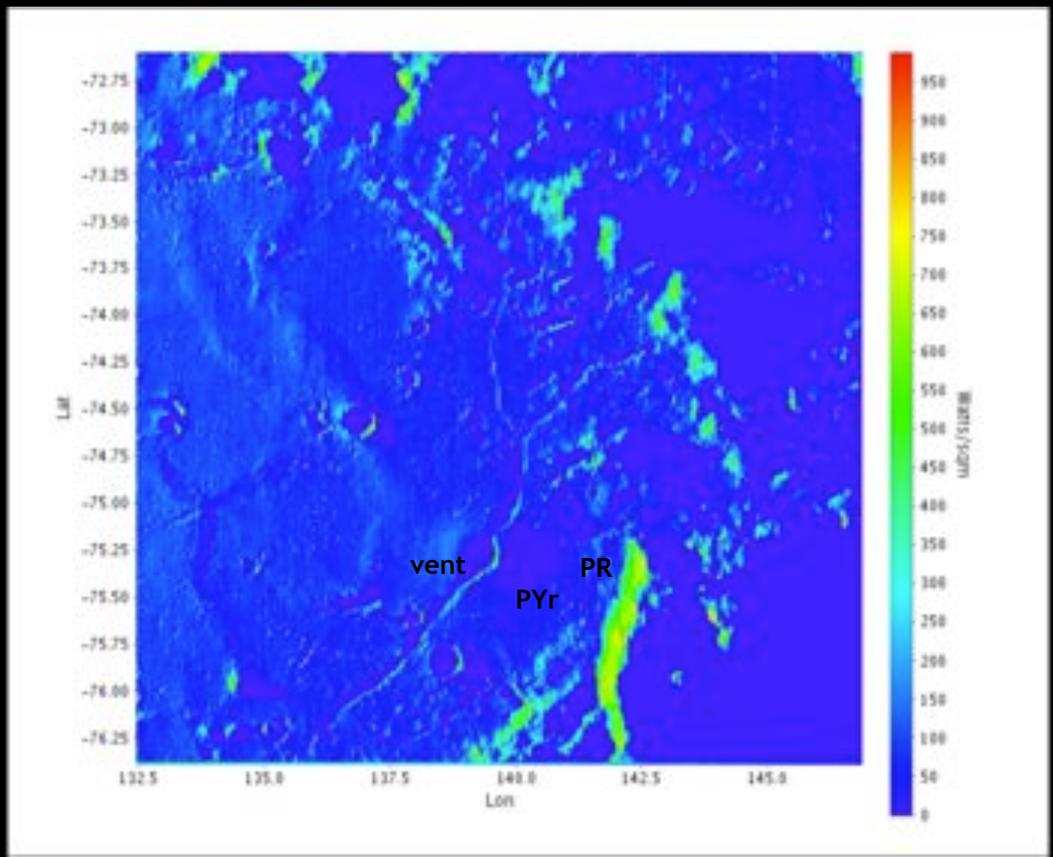
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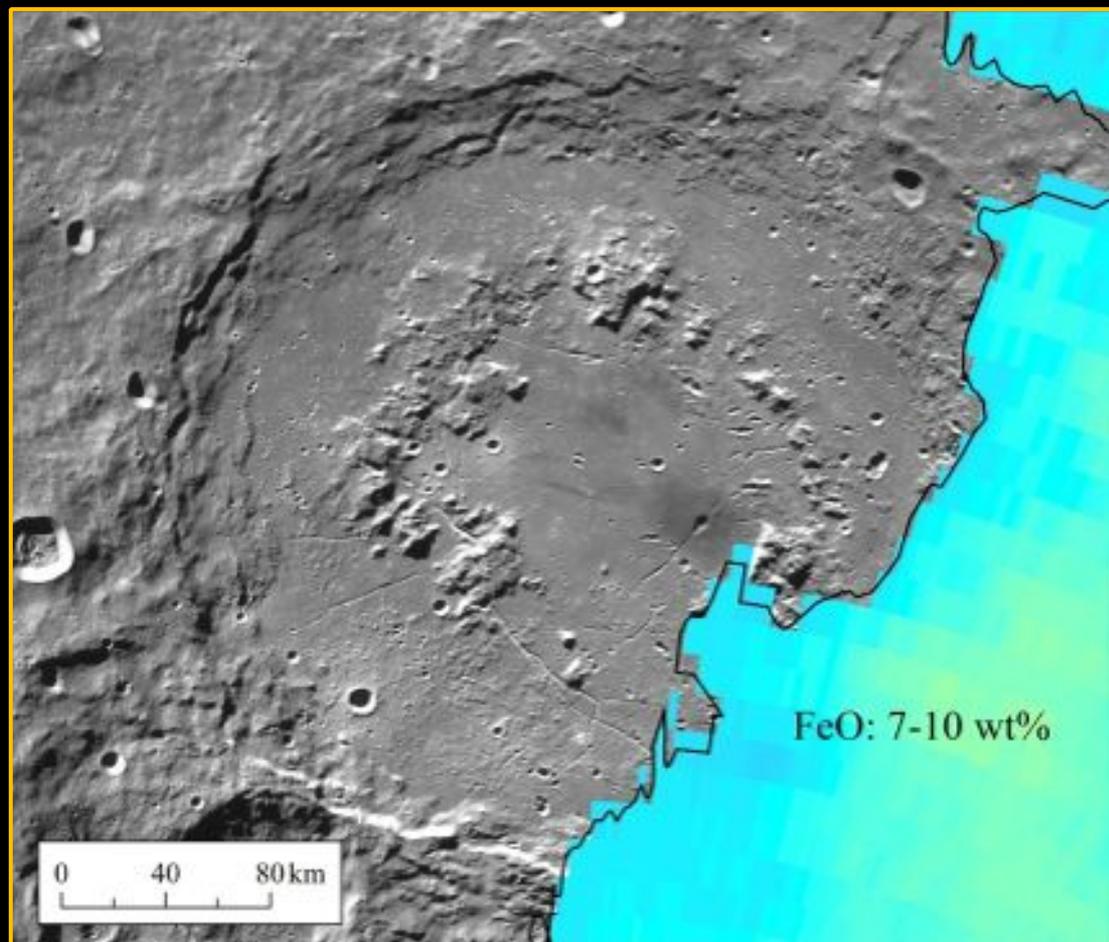
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Incidence  
sunlight  
(i.e., sunlight  
power on  
surface)



## POTENTIAL SCHRÖDINGER & SPA IMPACT MELT DEPOSITS

- Iron anomaly that has been used to define regions with SPA melt extends into the Schrodinger basin
- Within that region, look for low-Ca pyroxene exposures, which can reflect crystallized SPA melt



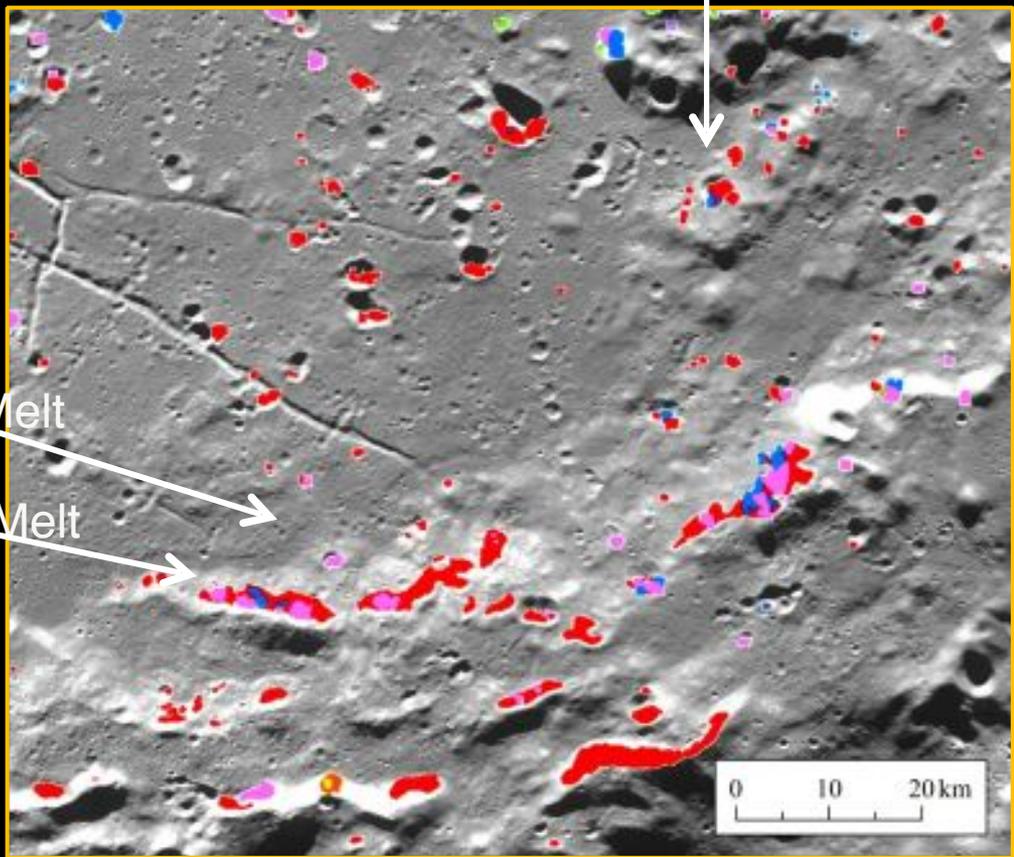
## POTENTIAL SCHRÖDINGER & SPA IMPACT MELT DEPOSITS

- Basin walls more likely to host SPA melt
  - Possible exposures of SPA melt
  - Slumped terraces, fallen rocks
- Possible Mission
  - Sample candidate SPA impact melt
  - Compare with Schrödinger melt

Candidate SPA Impact Melt

Schrödinger Impact Melt

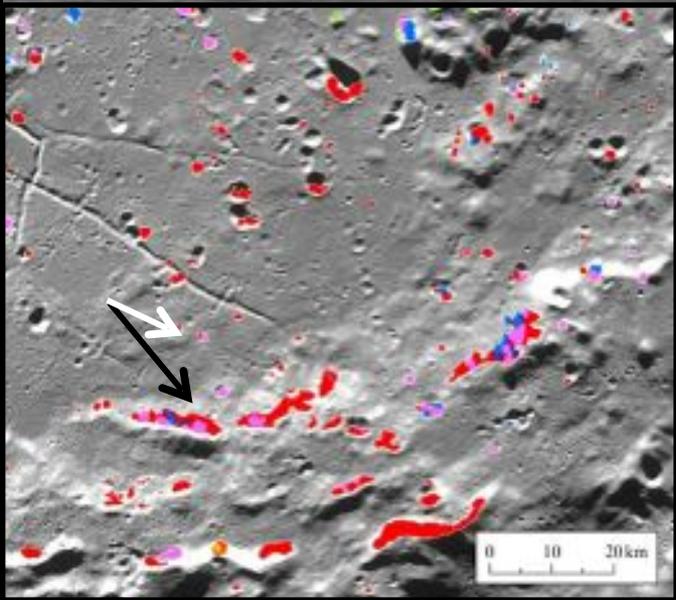
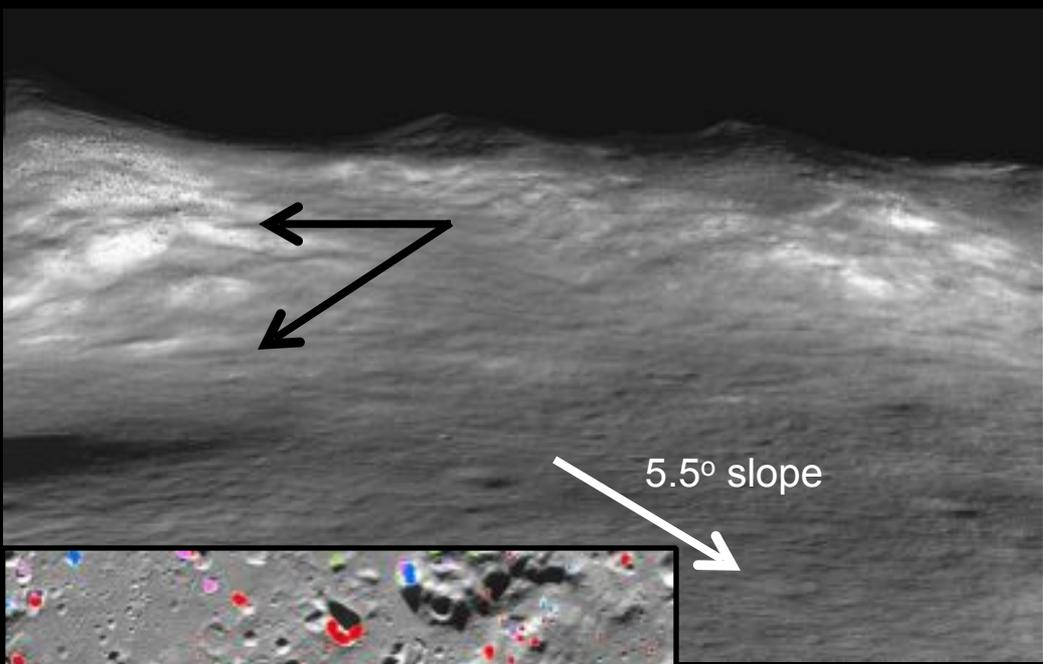
Candidate SPA Impact Melt



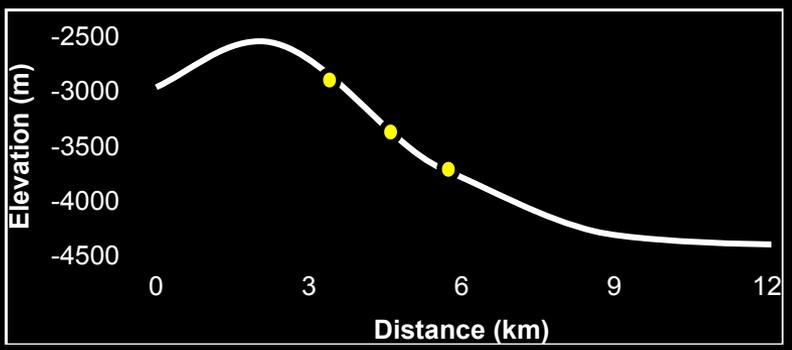
Low-Ca pyroxenes: red and pink  
Anorthosite (blue)  
Olivine (green)  
(Kramer et al. 2013)

Hurwitz & Kring (2014)

## POTENTIAL SCHRÖDINGER & SPA IMPACT MELT DEPOSITS

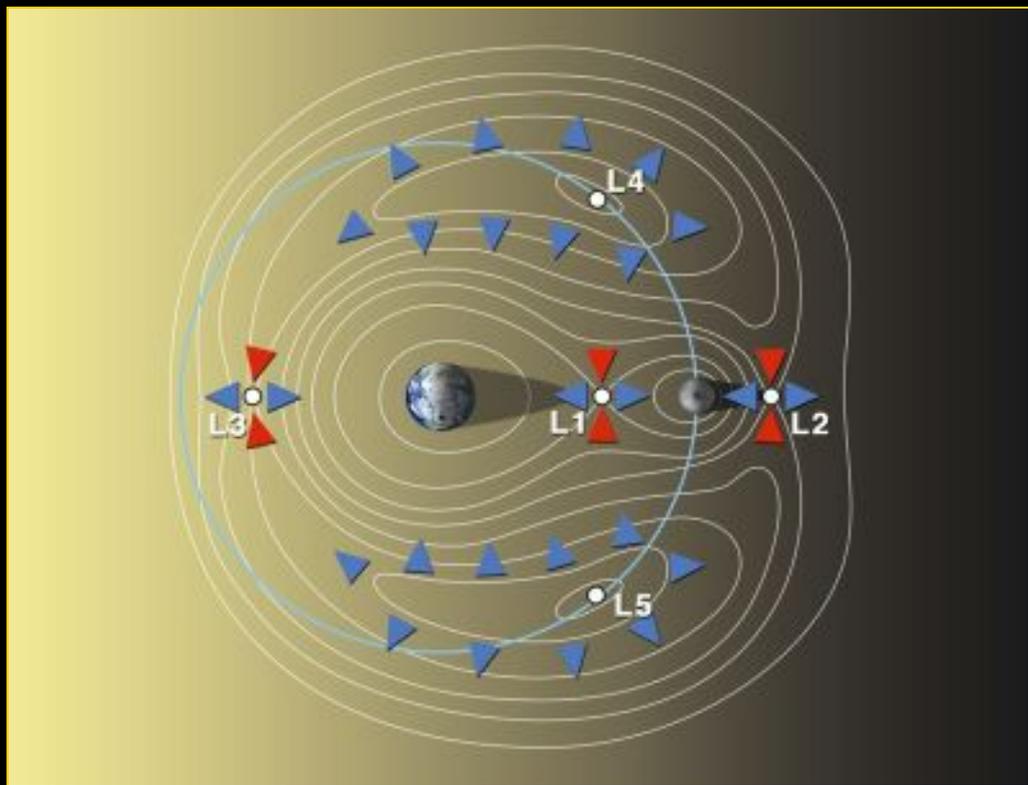


- Basin walls more likely to host SPA melt
  - Possible exposures of SPA melt
  - Slumped terraces, fallen rocks
- Possible Mission
  - Sample candidate SPA impact melt
  - Compare with Schrödinger melt
    - 15–18 km away on plains floor
    - Slope from plains to lower rocks: 5.5°



LOLA topography: 2x vertical exaggeration

## A SCIENCE PERSPECTIVE ABOUT HUMAN AND ROBOTIC EXPLORATION



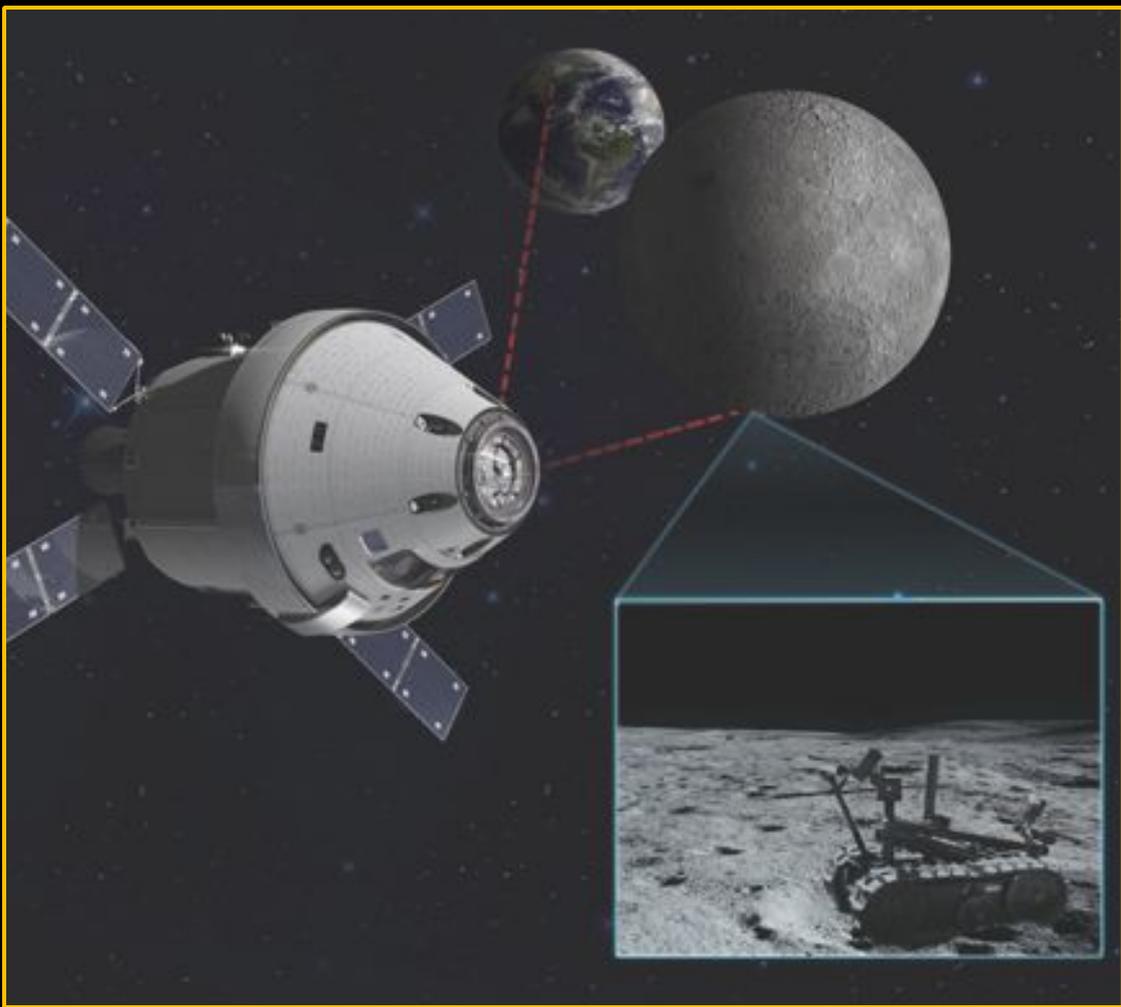
Burns, Kring, Norris, Hopkins, Lazio, & Kasper (2013)

An option we have been exploring

### Earth-Moon L2 Mission:

- L2 located 60,000 km above the lunar surface
- Orion launched and maneuvered into a halo orbit around L2
- The mission can also be conducted using the DRO architecture

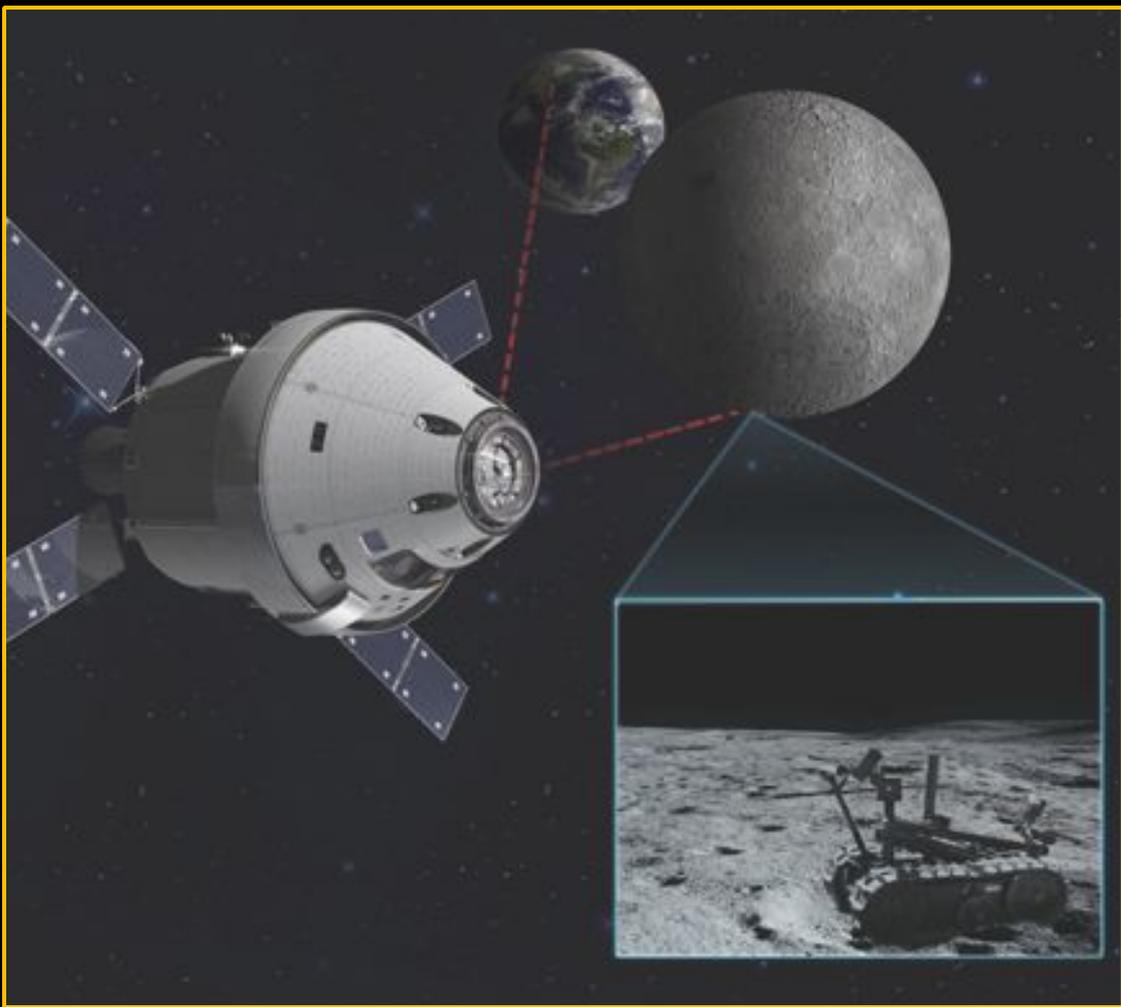
## PREPARING FOR HUMAN ASSISTED SAMPLE RETURN (PER THE GER)



### Exploration risk reduction:

- Demonstrate Orion in deep space and high speed Earth-entry
- 30 to 35 day mission into trans-lunar space
- Crew will travel 15% farther than Apollo and spend 3 times longer in deep space
- Practice tele-operation of rovers

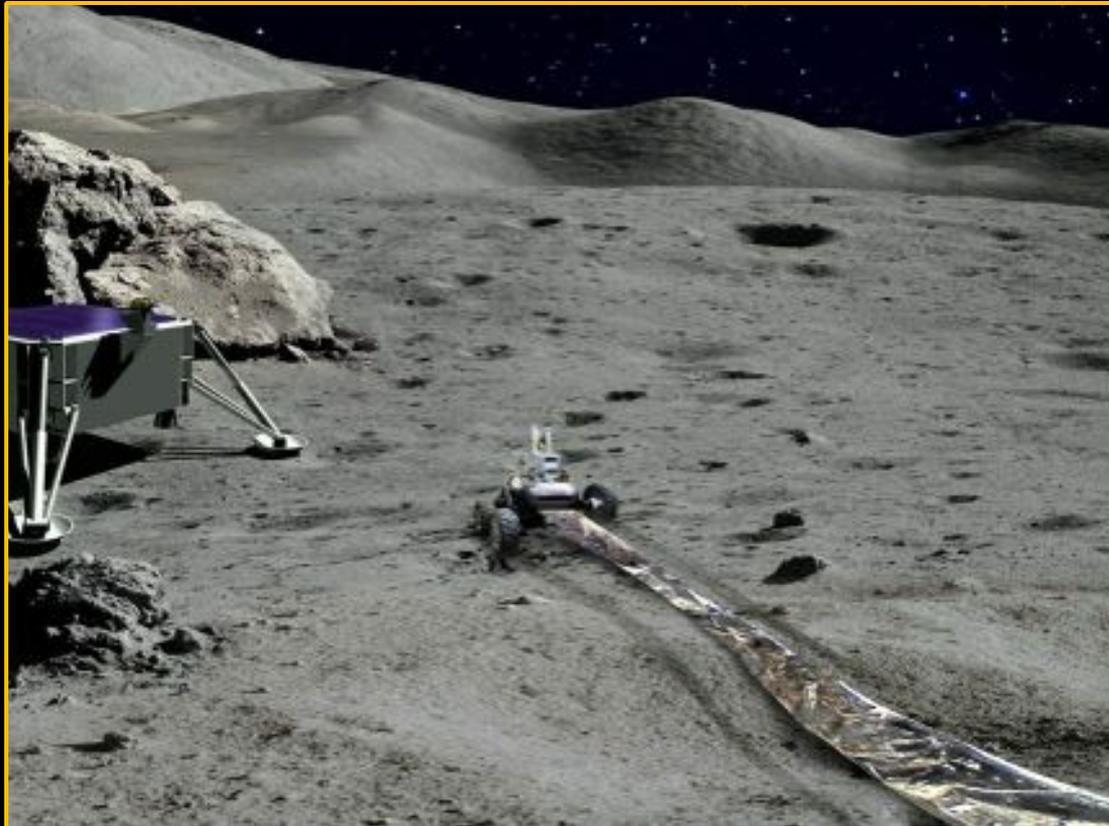
## PREPARING FOR HUMAN ASSISTED SAMPLE RETURN (PER THE GER)



### Science objectives:

- Land and explore a region within SPA (**for example, Schrödinger Basin**)
- **Geologic measurements will be made.**
- **A sample will be collected and returned to Earth.**
- **An astrophysical system will be deployed.**

**Status: Integrated science and engineering studies continue.**



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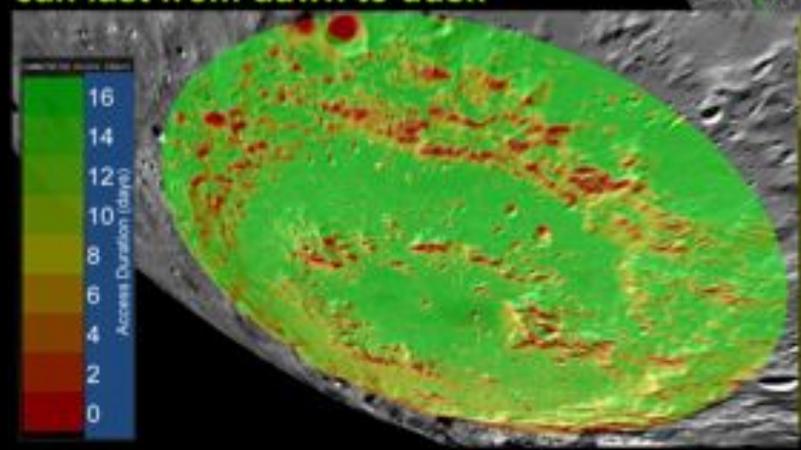
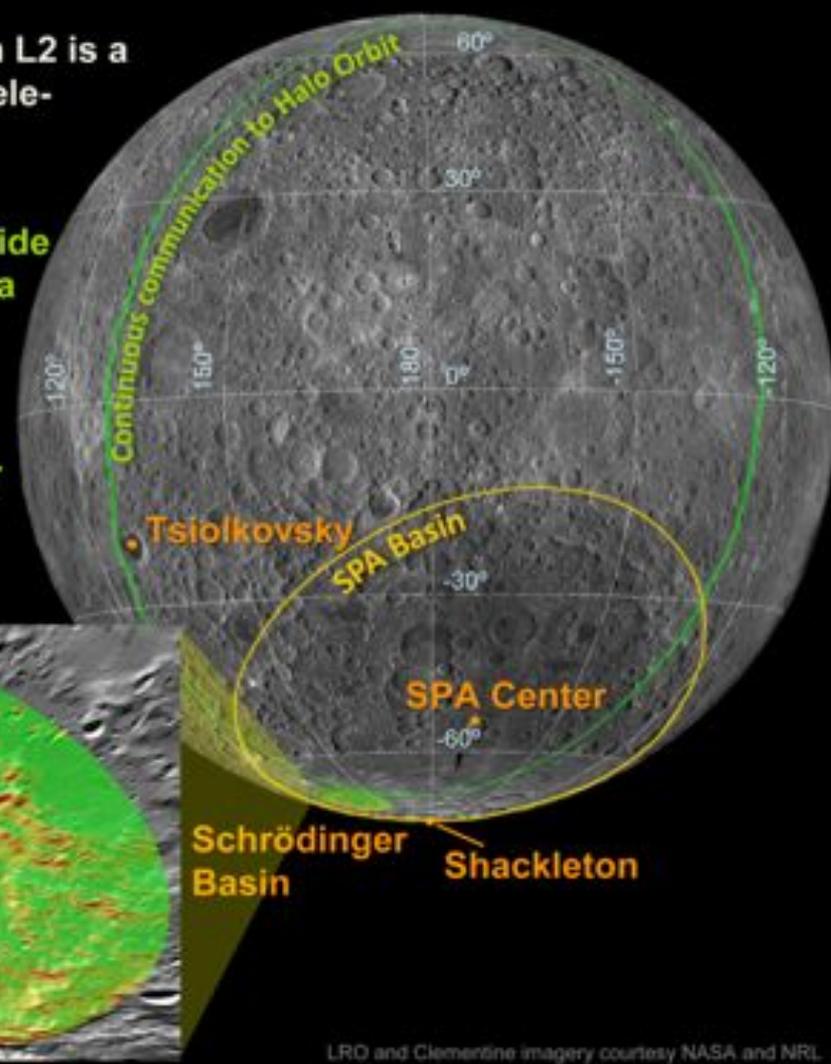
## COMMUNICATION COVERAGE FROM A HALO ORBIT ABOUT EM-L2

Lockheed Martin (courtesy of Josh Hopkins)

An Orion mission to Earth-Moon L2 is a feasible alternative to DRO for tele-robotics missions

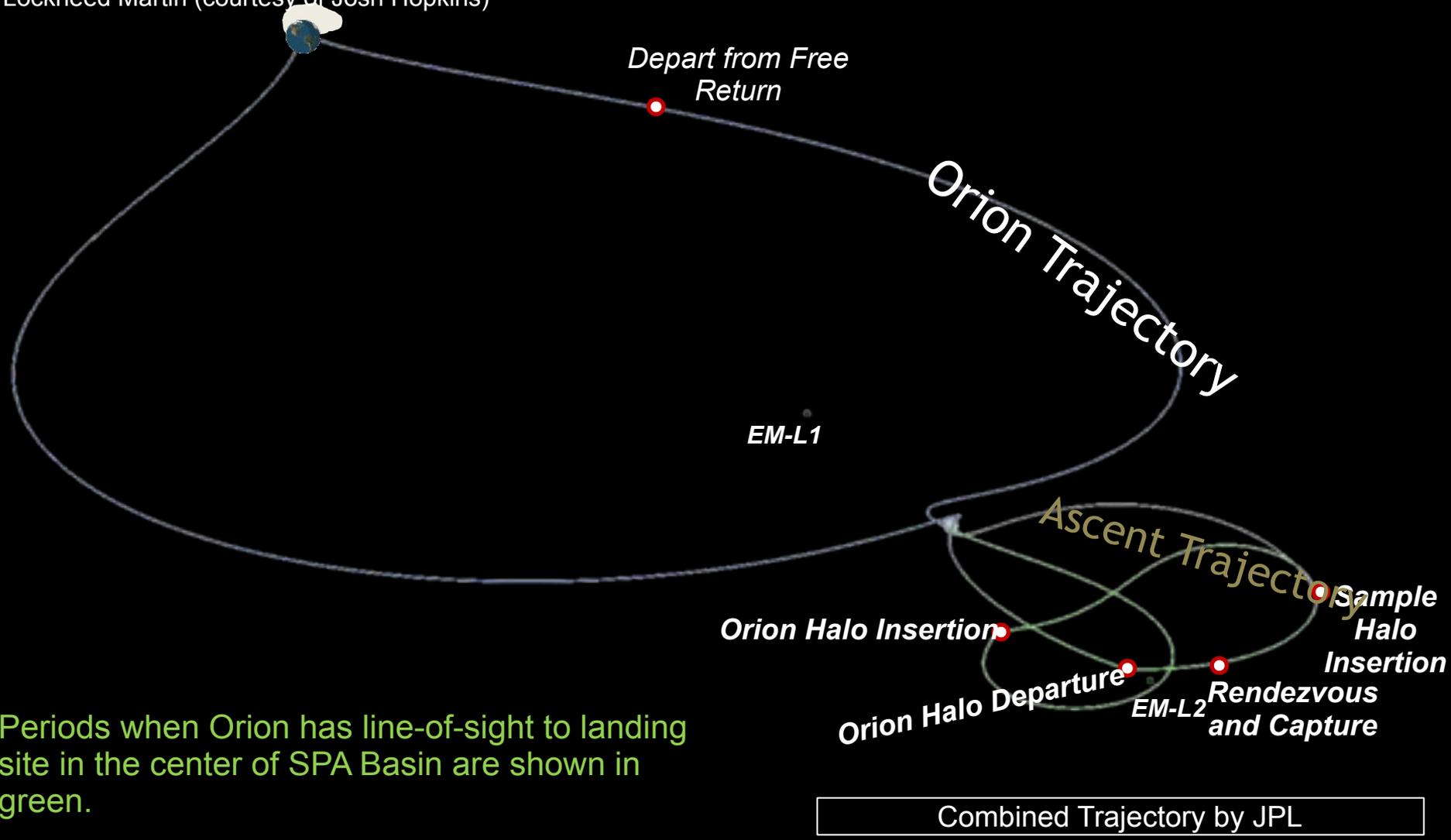
A large area of the lunar farside is continuously visible from a compact L2 halo orbit

Even in limb sites like Schrödinger Basin there are areas where communication to the halo orbit can last from dawn to dusk



## TRAJECTORY FOR SAMPLE CAPTURE IN EARTH-MOON L2

Lockheed Martin (courtesy of Josh Hopkins)



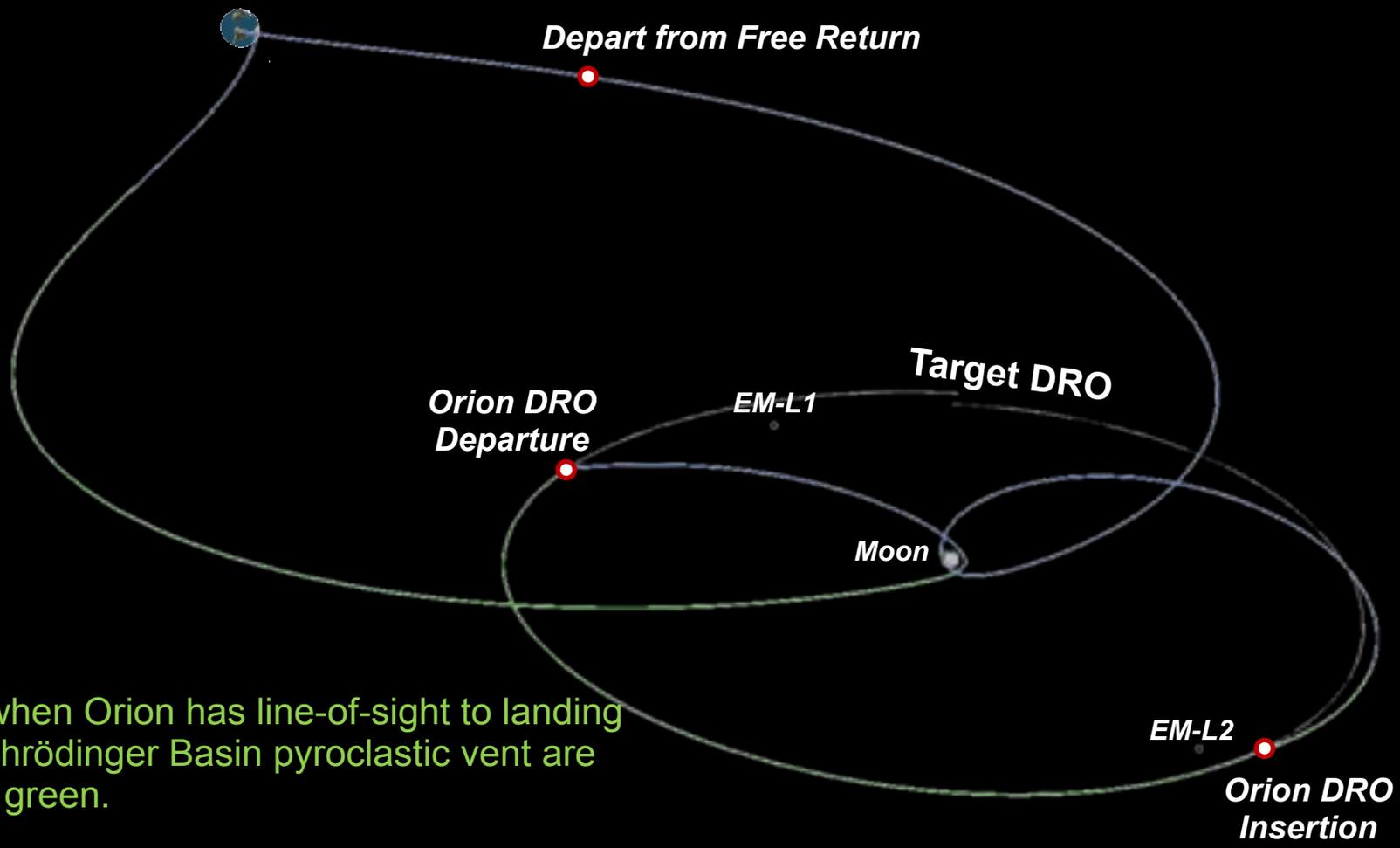
Periods when Orion has line-of-sight to landing site in the center of SPA Basin are shown in green.

Combined Trajectory by JPL



## TRAJECTORY FOR TELEROBOTICS IN A DISTANT RETROGRADE ORBIT (DRO)

Lockheed Martin (courtesy of Josh Hopkins)

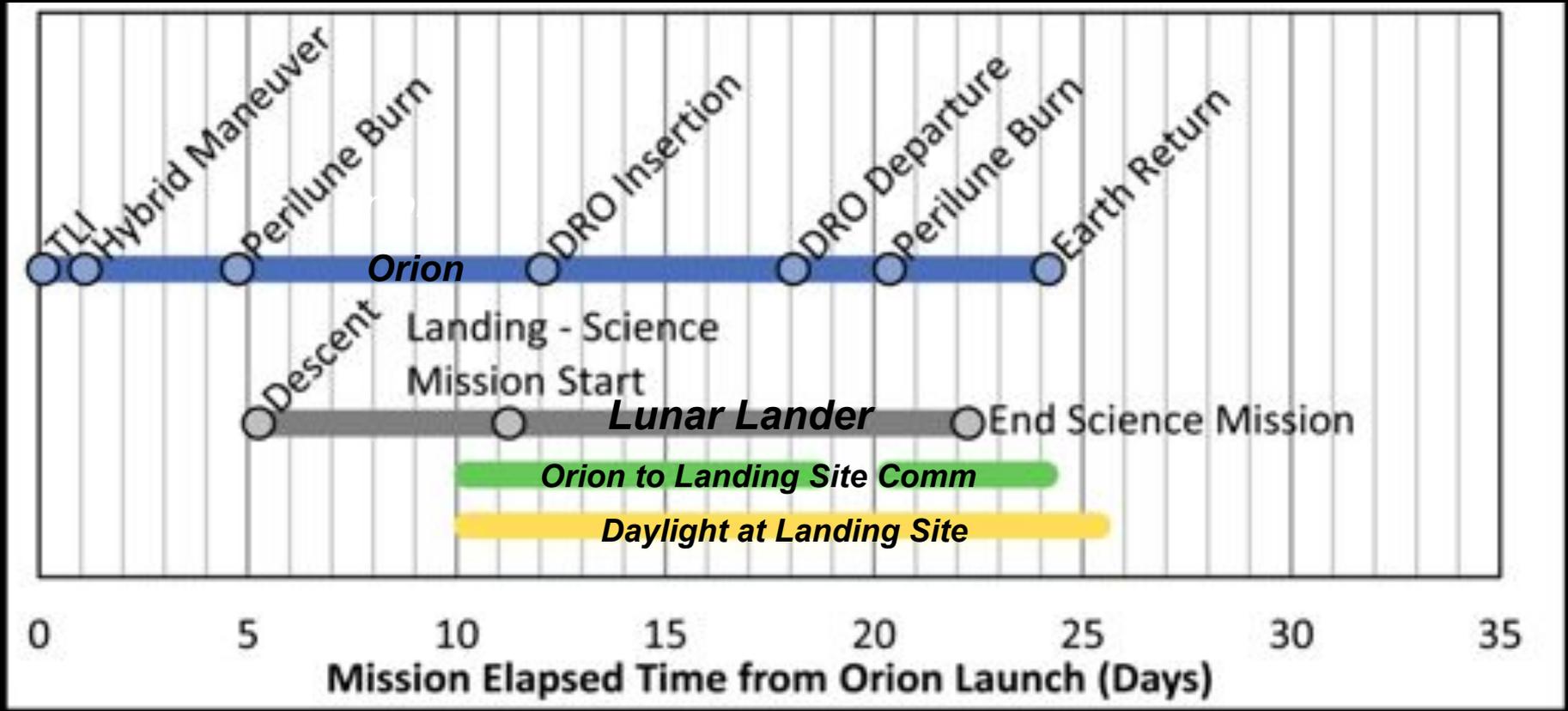


Periods when Orion has line-of-sight to landing site at Schrödinger Basin pyroclastic vent are shown in green.

Trajectory is compatible with Orion delta-V and duration capabilities

## TIMELINE FOR DRO/SCHRÖDINGER BASIN MISSION

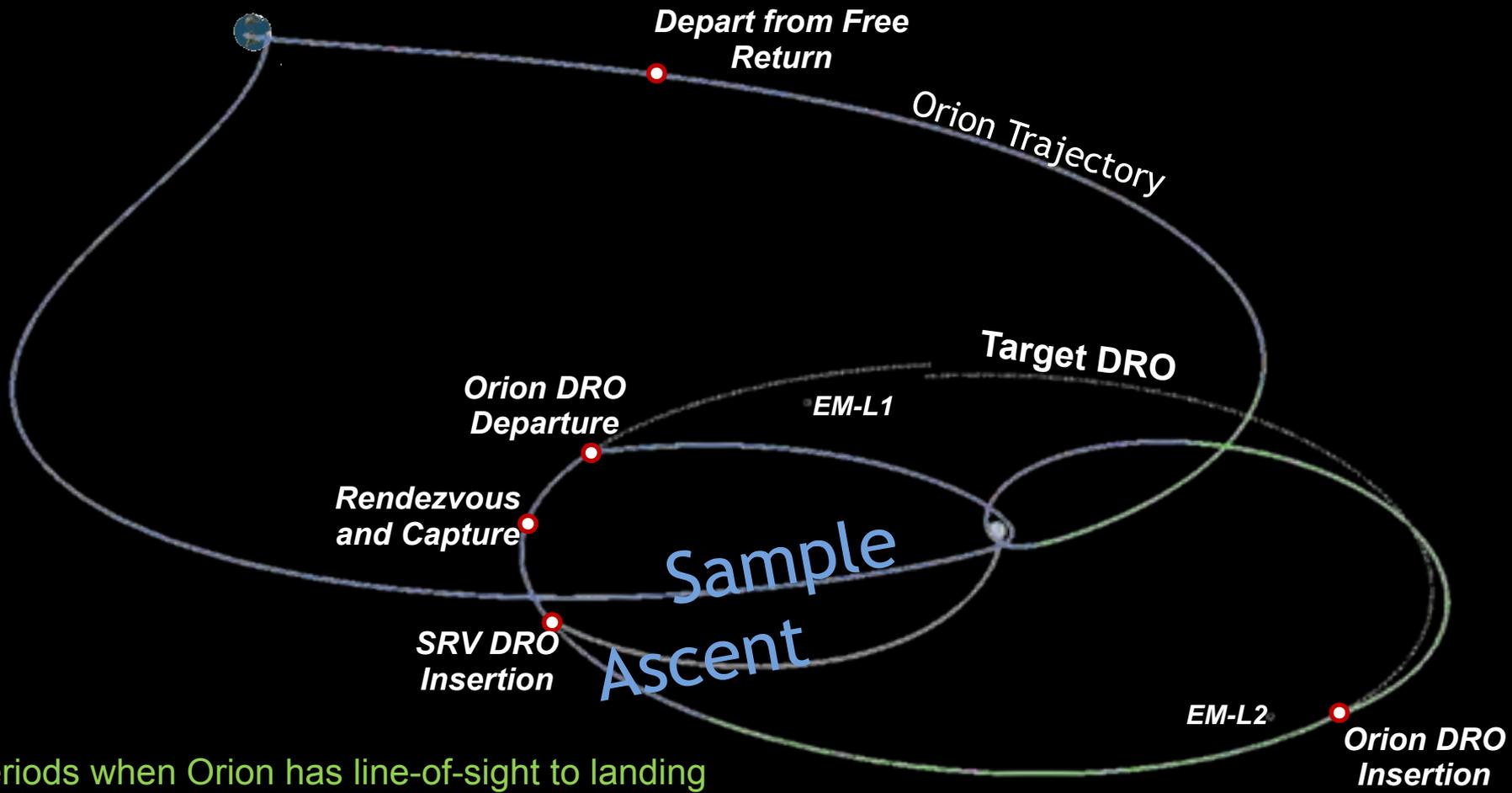
Lockheed Martin (courtesy of Josh Hopkins)



Telerobotic rover has 11 day operating duration with one telecom gap of ~1.5 days when Orion is not visible.

## TRAJECTORY FOR SAMPLE CAPTURE IN DRO FROM SPA BASIN

Lockheed Martin (courtesy of Josh Hopkins)



Periods when Orion has line-of-sight to landing site in the center of SPA Basin are shown in green.

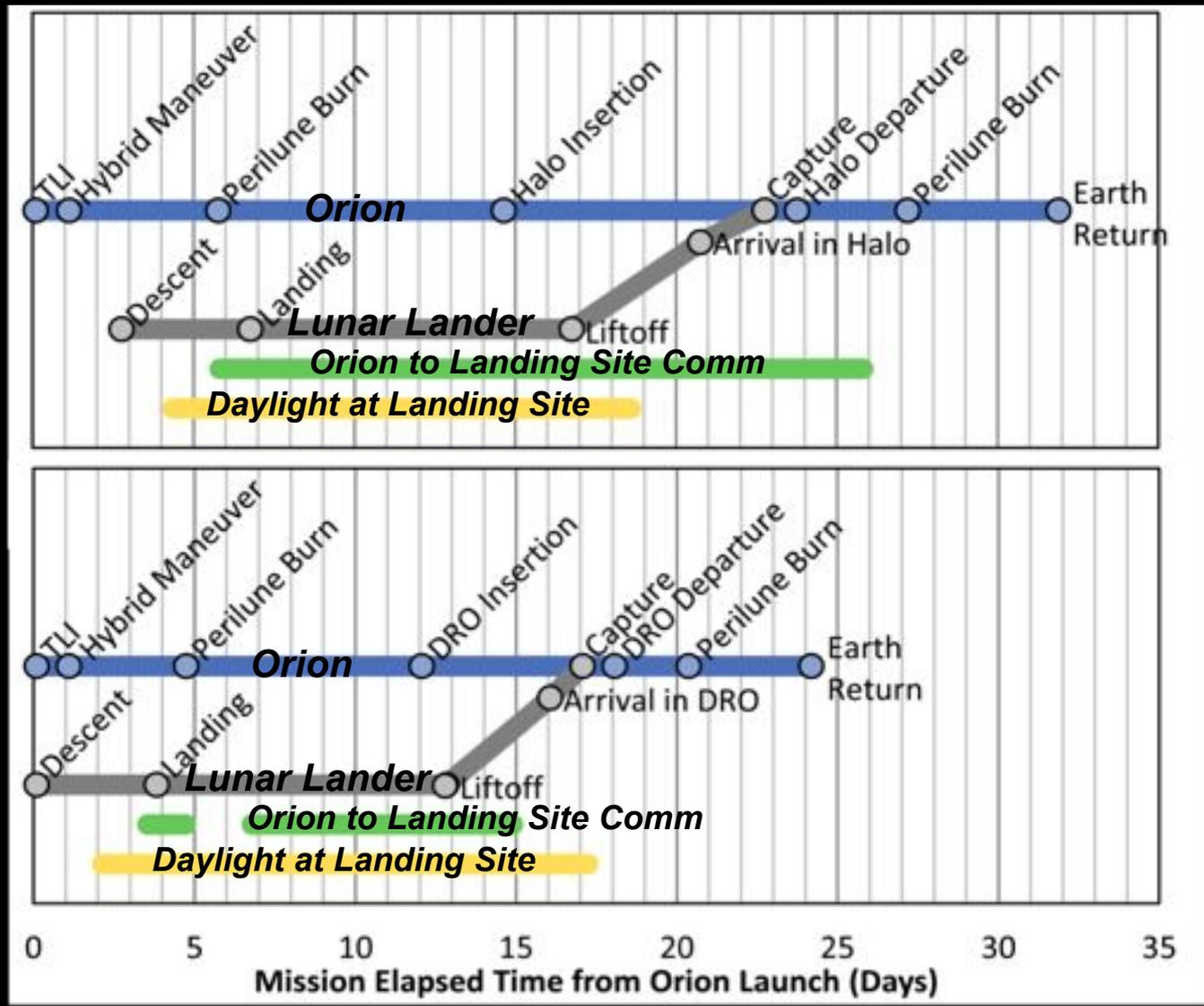


## SAMPLE CAPTURE MISSION TIMELINE COMPARISON

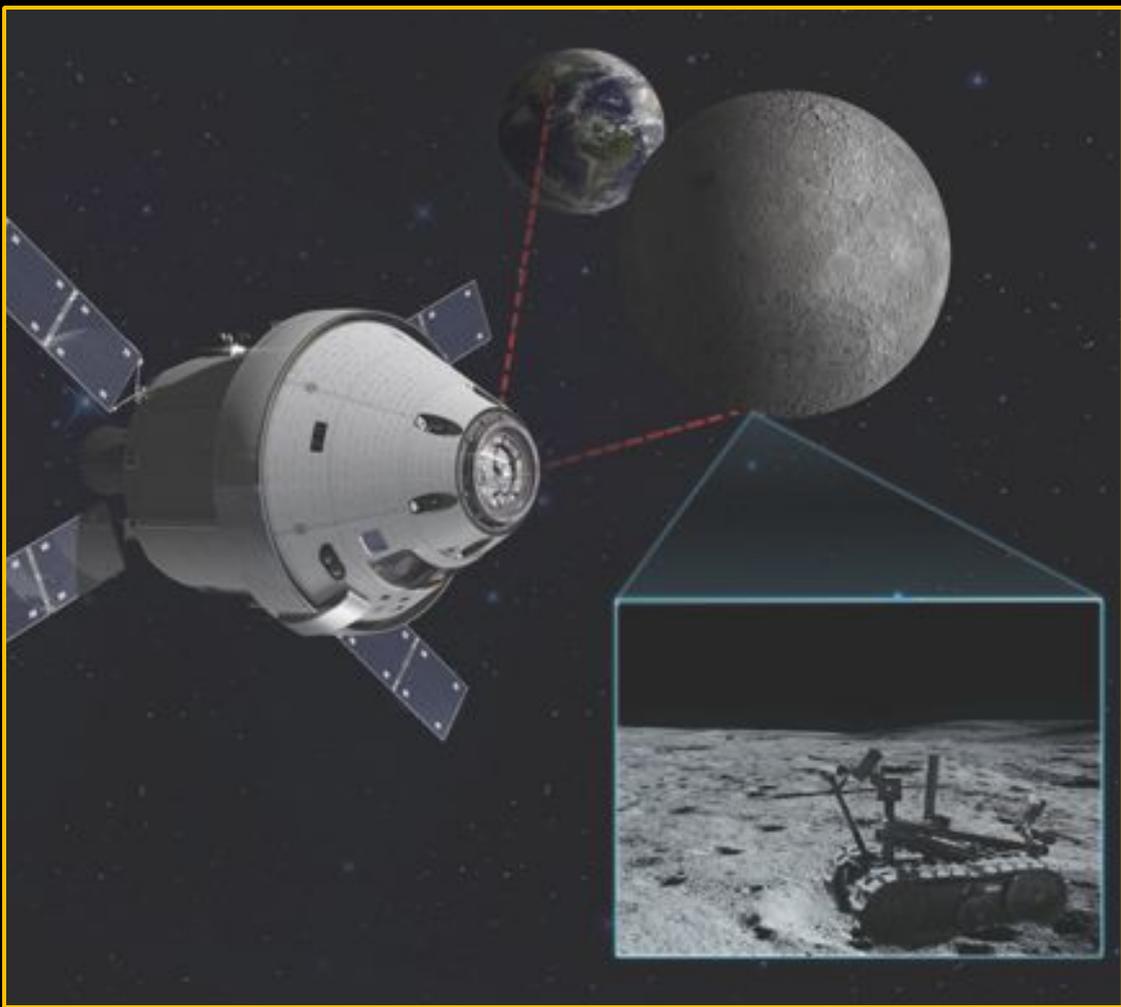
Lockheed Martin (courtesy of Josh Hopkins)

Sample Return Central SPA Basin to L2

Sample Return Central SPA Basin to DRO

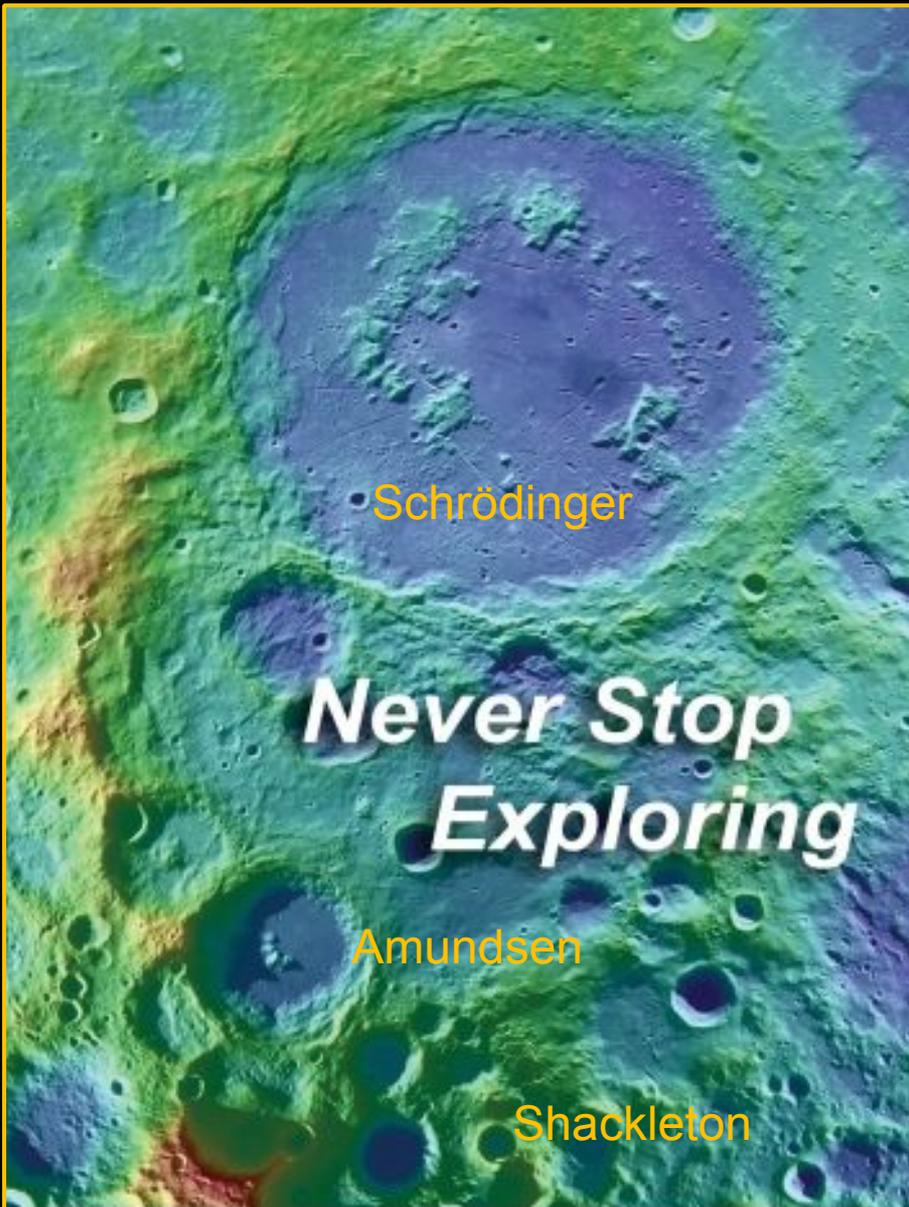


## PREPARING FOR HUMAN ASSISTED SAMPLE RETURN (PER THE GER)



### Additional options:

- Deploy a communication satellite from Orion to support additional surface activity after crew returns to Earth
- If long-term station-keeping by crew is implemented in an L2 or distant retrograde orbit (DRO), then additional tele-ops can be conducted with the first rover and potentially other landed assets.



## Conclusions

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Schrödinger basin is one of the highest priority landing sites based on a global assessment of the NRC (2007) objectives

It can be used to test the

- Lunar Cataclysm Hypothesis
- Lunar Magma Ocean Hypothesis

It can provide ISRU resources

- Pyroclastic deposits
- & potentially volatile-rich deposits

The Moon is the best and most accessible place in the Solar System to answer fundamental planetary science questions.